

# Calculating and Applying BSDFs in *Radiance*

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Anyhere Software



# Goals

- Support WINDOW 6 BSDF output (XML)
- Create WINDOW 6 BSDFs from geometric models using freeware (i.e., *Radiance*)
- Render accurate images with direct shading using **mkillum**
- Perform efficient annual simulations
- Support higher-resolution BSDF data

# Status

- ❧ WINDOW 3 BSDF format still subject to revision
- ❧ Support for reading and creating BSDFs is currently limited to transmittance data
- ❧ **mkillum** has been tested and refined to produce reasonable accuracy for diffusing systems
- ❧ New **genBSDF** tool for creating BSDF files from geometric (*Radiance* or MGF) descriptions
- ❧ Annual simulation for the **daring** with **rtcontrib**



# What Is a BSDF?

- “BSDF” stands for “Bidirectional Scattering Distribution Function” (BRDF + BTDF)
- Describes how light scatters off a surface
- Incl. wavelength, a 5-dimensional scalar function
  - unitless ratio of outgoing radiance over incoming irradiance (1/steradian)

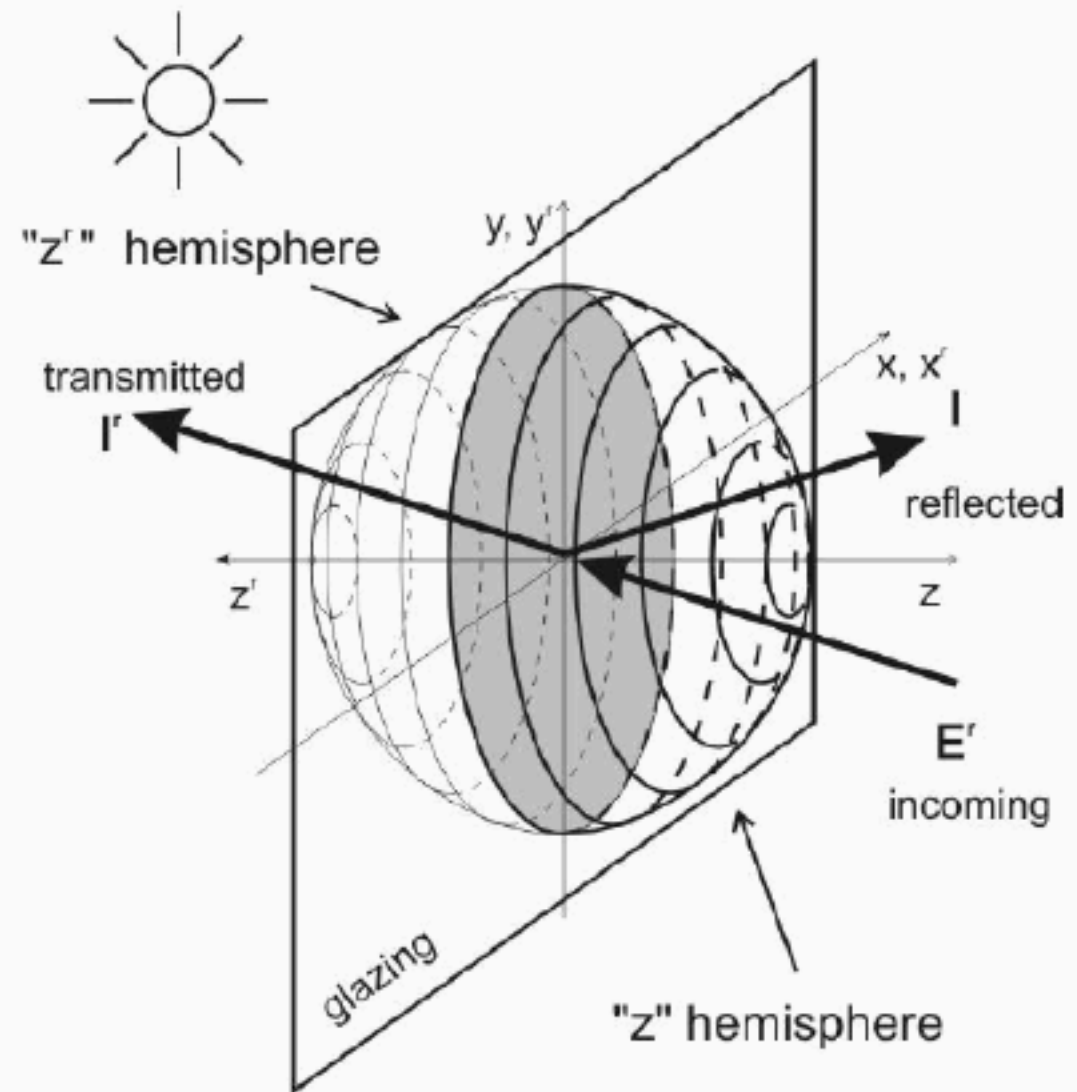
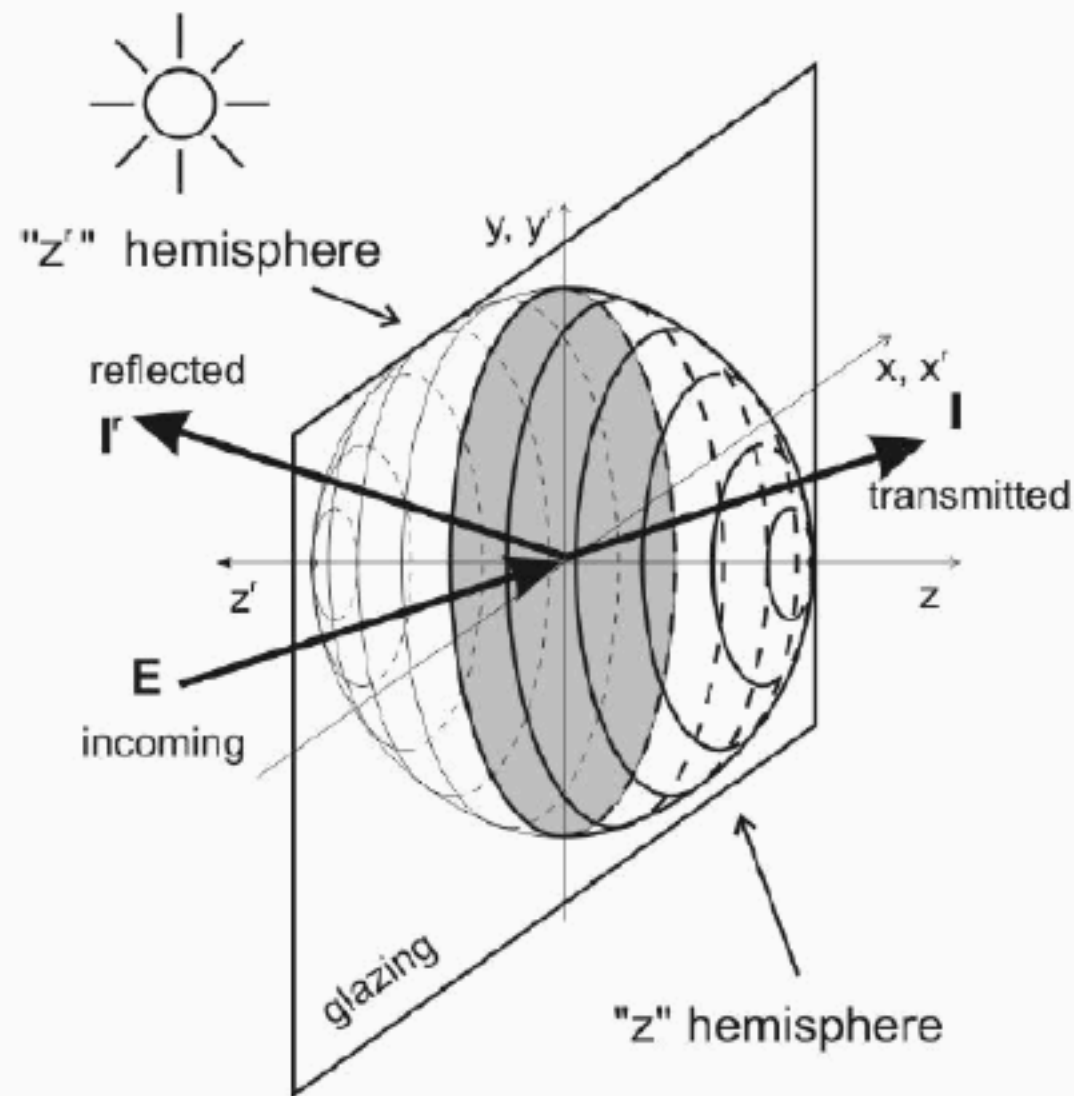
# WINDOW 6 BSDF

- ❧ Breaks BSDF into integrated spectra:
  - ❧ Visible (380-780 nm), Solar (300-2500 nm), NIR (780-2500 nm), FIR (5000- 40000 nm)
- ❧ Breaks scattering into four components:
  - ❧ Front Transmission, Back Transmission, Front Reflection, Back Reflection

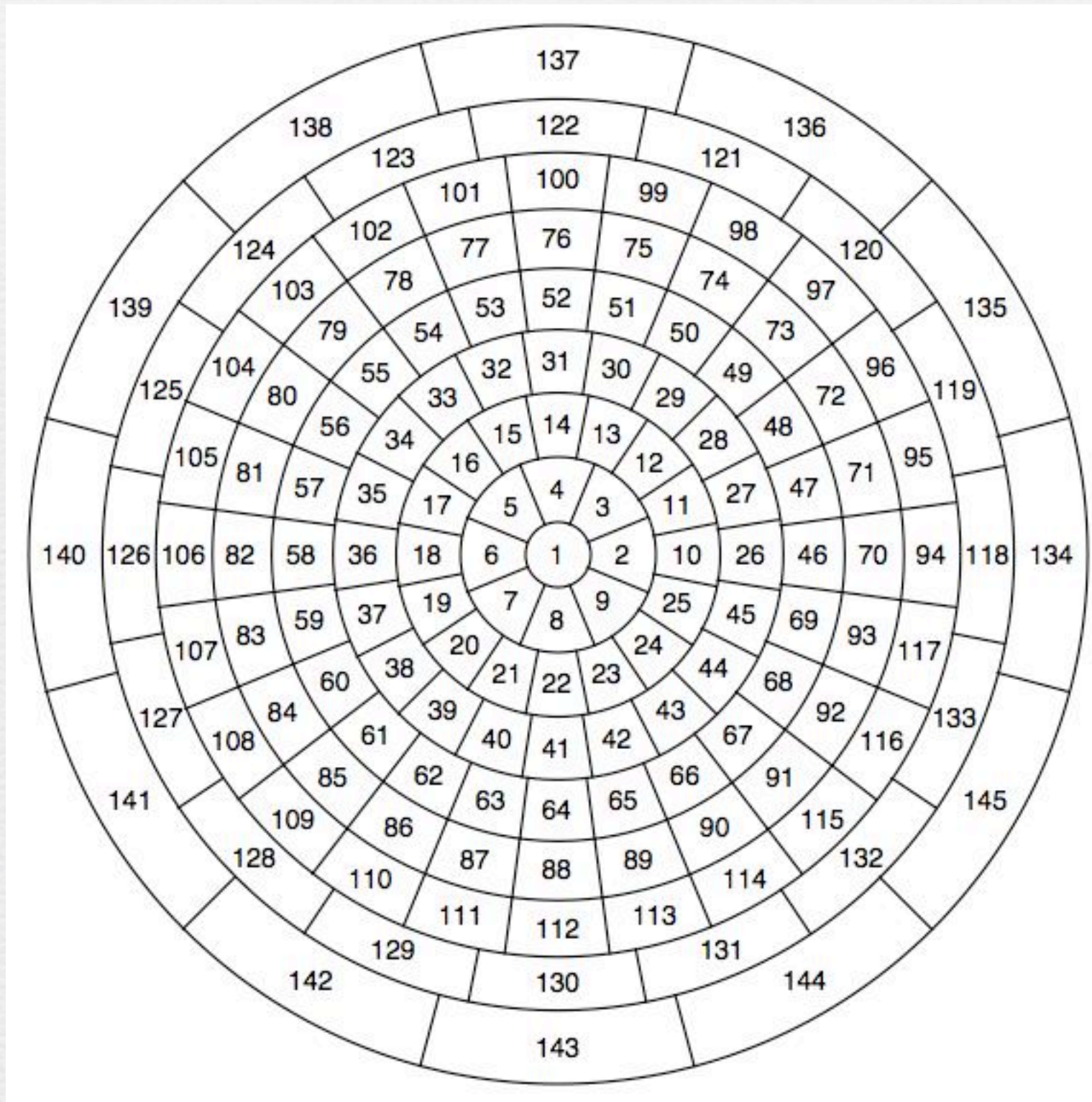


# BSDF Coordinates (1)

## Klems/Window 6 system



# BSDF Coordinates (2)





# WINDOW 6 XML File

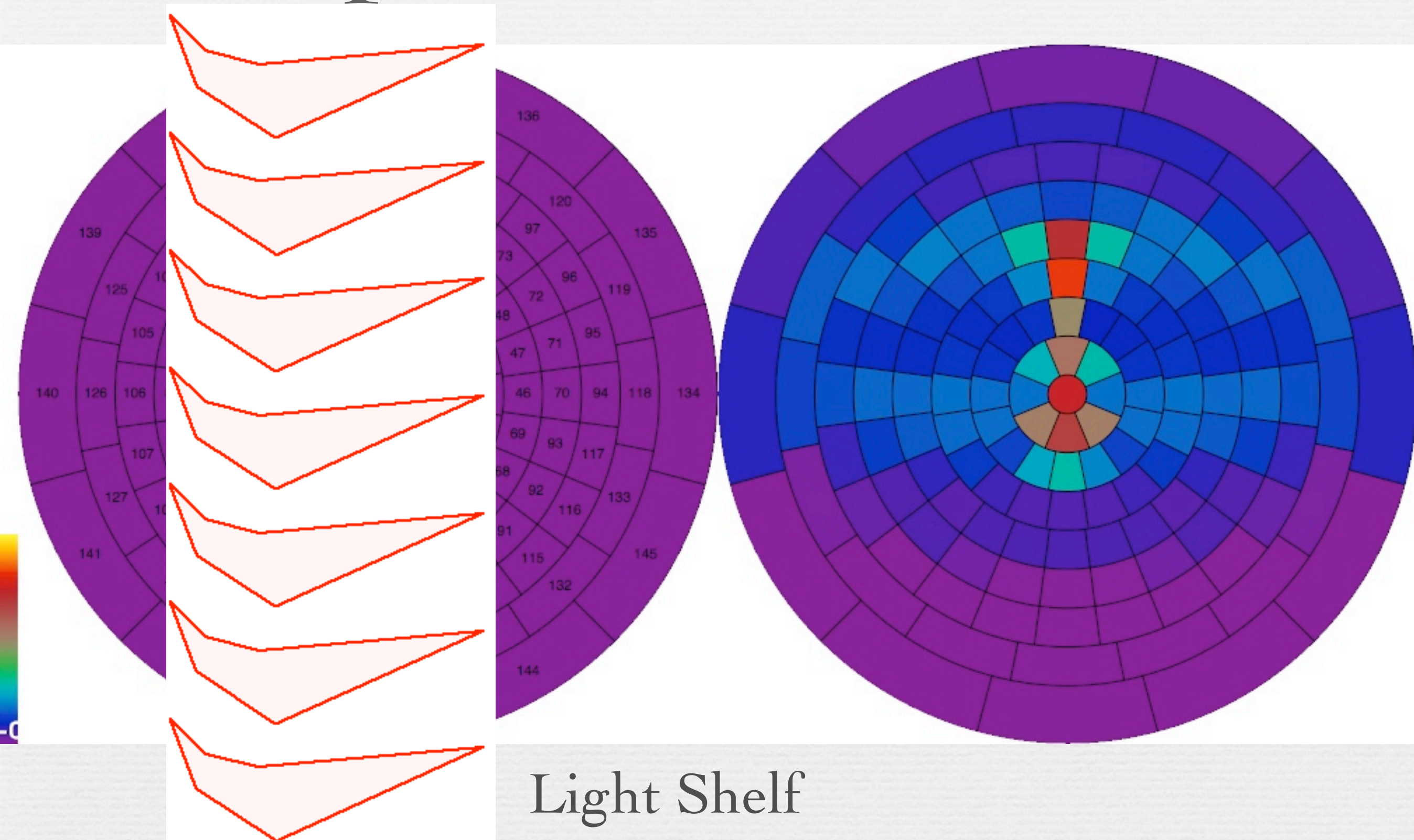
```
<WavelengthData>
<Wavelength unit="Integral">NIR</Wavelength>
<SourceSpectrum>CIE Illuminant D65 1nm.ssp</SourceSpectrum>
<DetectorSpectrum>ASTM E308 1931 Y.dsp</DetectorSpectrum>
<WavelengthDataBlock>
  <WavelengthDataDirection>Transmission Front</WavelengthDataDirection>
  <ColumnAngleBasis>LBNL/Klems Full</ColumnAngleBasis>
  <RowAngleBasis>LBNL/Klems Full</RowAngleBasis>
  <ScatteringDataType>BTDF</ScatteringDataType>
  <ScatteringData>
2.443881,    0.047337,    0.041435,    0.038990,    0.041435,
0.047337,    0.048413,    0.046964,    0.048413,    0.047337,
0.040883,    0.035154,    0.031478,    0.030108,    0.031363,
0.035154,    0.040605,    0.047337,    0.048086,    0.044691,
0.042586,    0.042007,    0.042537,    0.044691,    0.047921,
0.047337,    0.038892,    0.031273,    0.025227,    0.021345,
0.020007,    0.021345,    0.025227,    0.031273,    0.038892,
...
```



# Getting BSDF Data

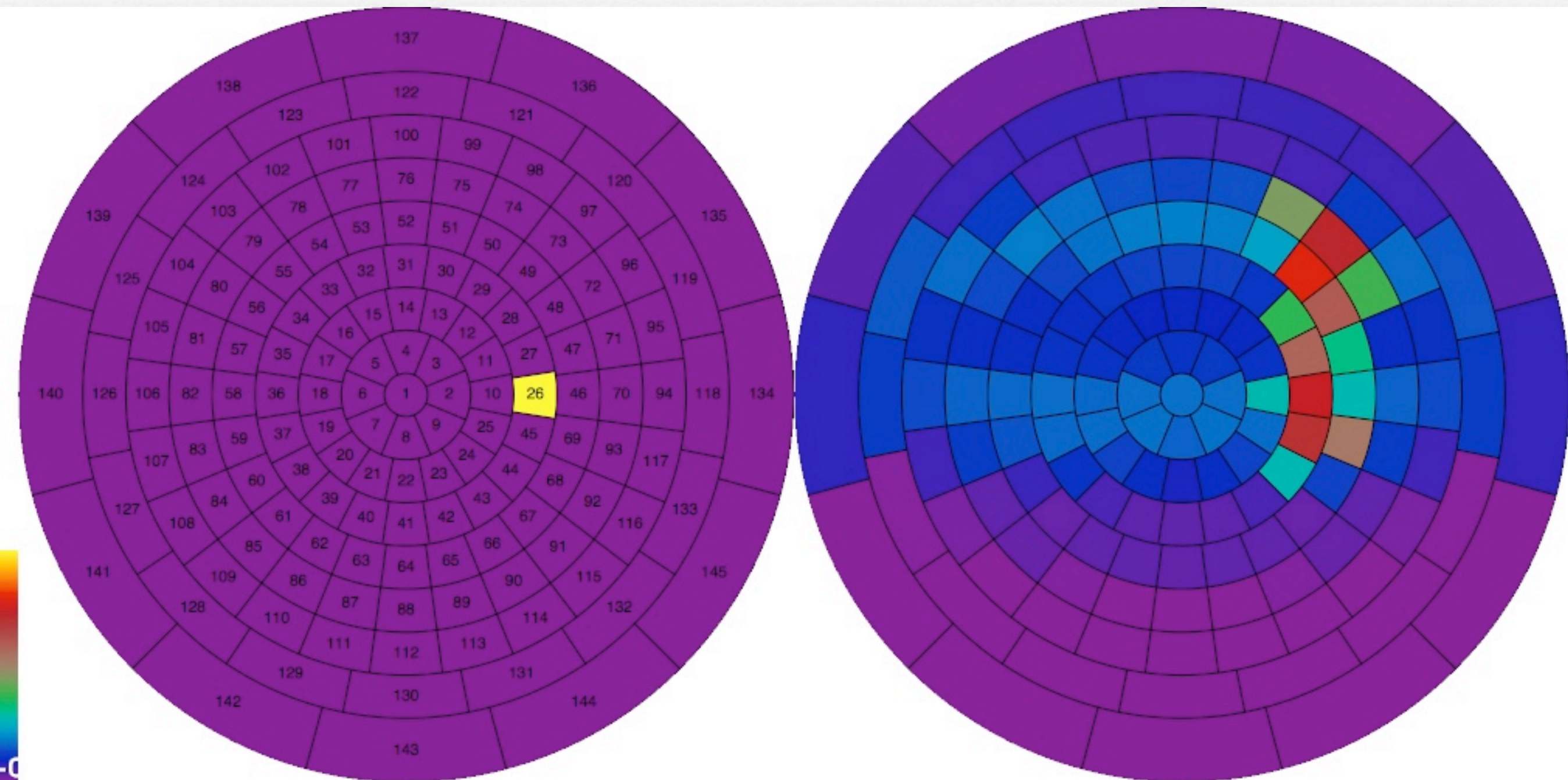
- ❧ WINDOW 6 will have access to a database of fenestration systems
- ❧ Systems may be combined in “layers” up to a point
- ❧ Some BSDFs are measured, most are simulated
  - ❧ BSDF measurements go into simulations, too
- ❧ New **genBSDF** program in *Radiance* 4.1

# Sample BTDF Data (1)





# Sample BTDF Data (2)



Light Shelf

Visualization by Andrew McNeil

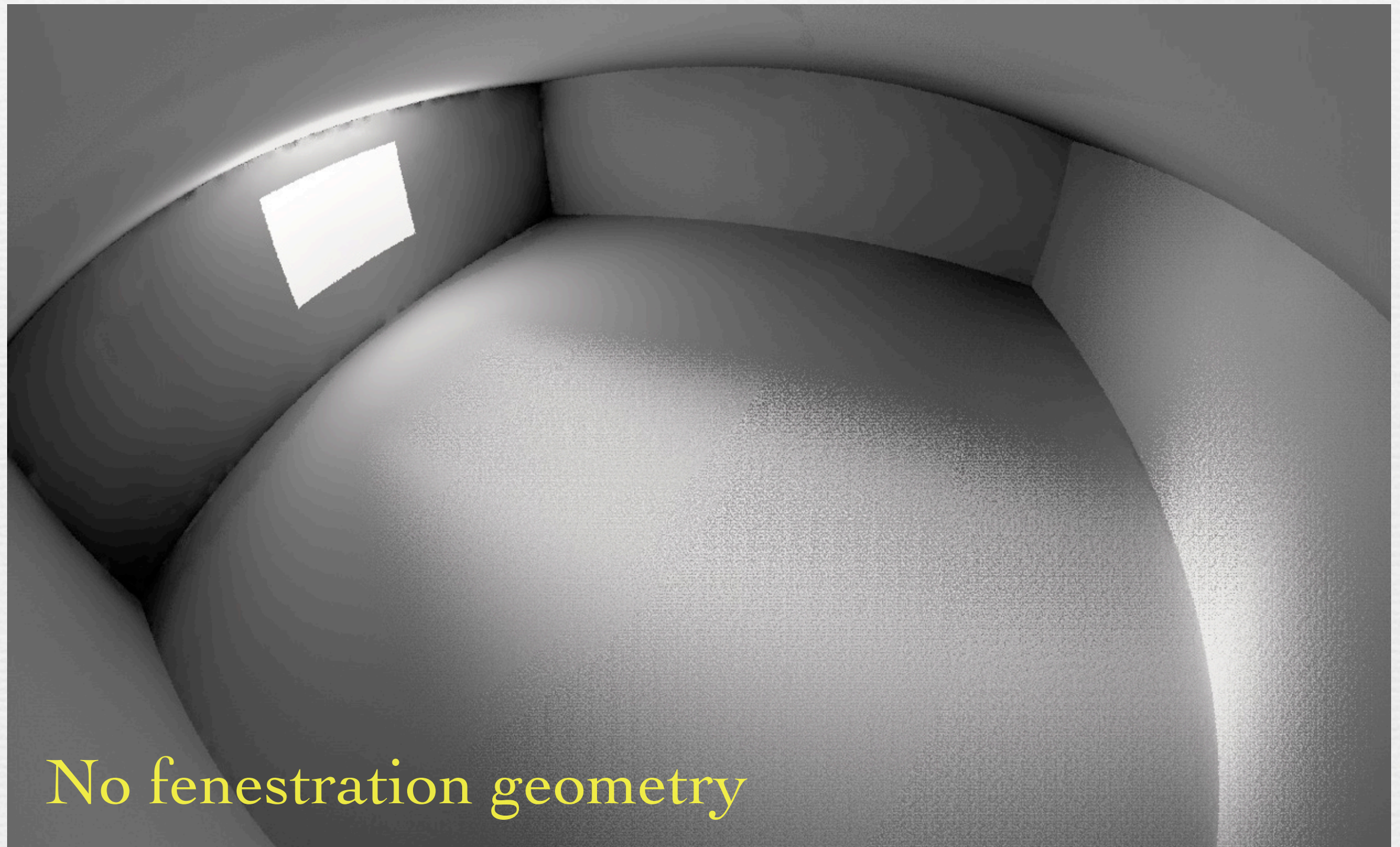


# BSDFs in **mkillum**

- Since 3.9 (2008), **mkillum** has supported BTDFs
  - fully debugged by 4.0 (Spring 2010)
- Works best when:
  - BTDF is somewhat diffuse, or
  - model includes geometry for direct component



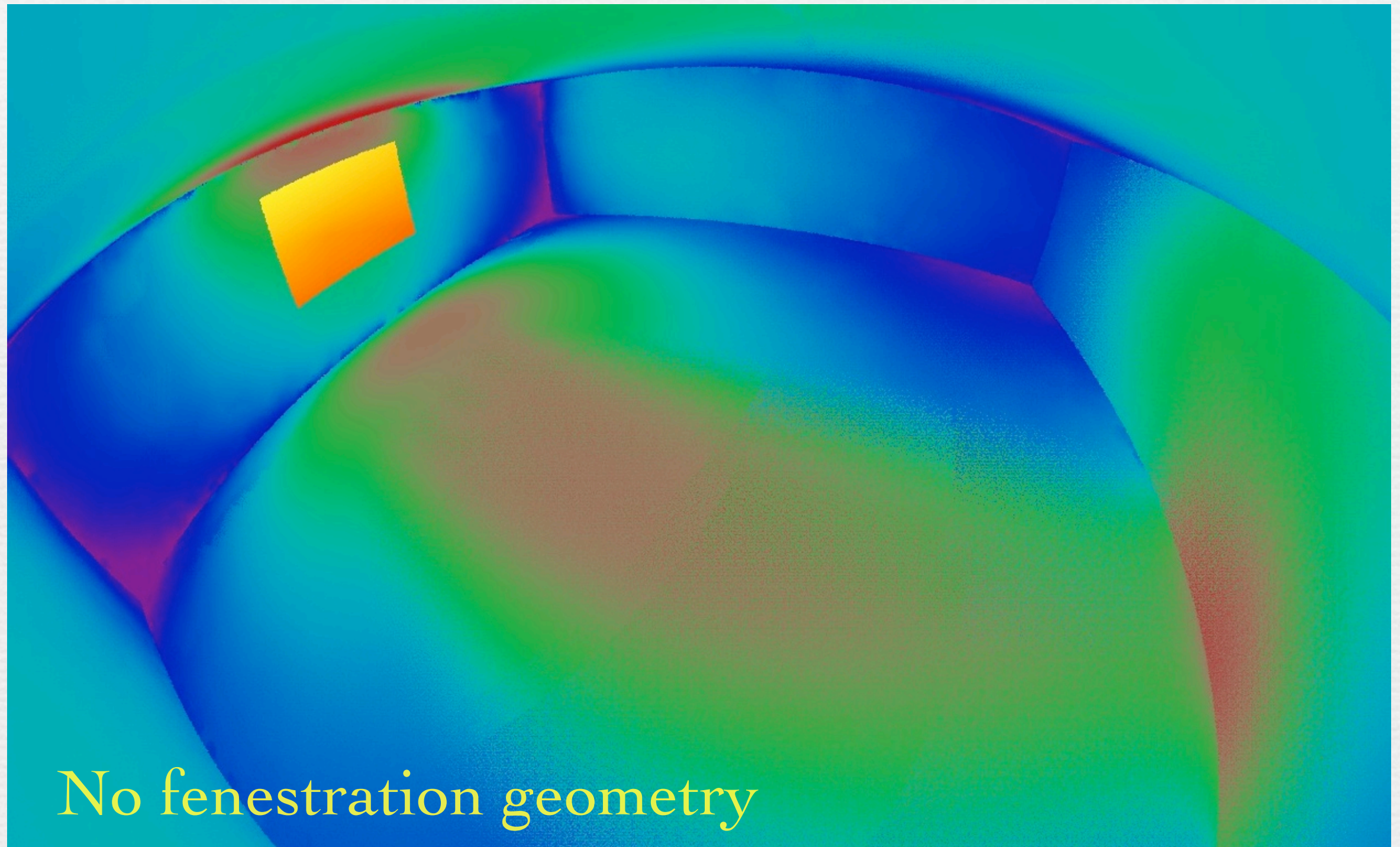
# Example Results



No fenestration geometry



# Averages Agree





# Specular Challenge

- 145 hemisphere directions cover 10-30° regions
  - The sun covers about 0.5°
- Breaking hemisphere further leads to drastic increases in computational load
  - 80000 divisions needed for sun-sized patches
  - that's an 80000x80000 BTDF (6.4e9 values)
  - ...and still no direct solar pattern on interior

# BSDF with Geometry

- ✧ Incorporate model of fenestration device in XML
- ✧ Use MGF (Materials and Geometry Format) developed for IESNA luminaire descriptions
  - ✧ simple, public <<http://radsite.lbl.gov/mgf>>
  - ✧ embeddable in XML
  - ✧ standard C parser makes support easy



# Sample MGF

```
<Geometry format="MGF" unit="Meter"> XML embedding
# Y-axis points "up", Z-axis into room, right-handed coordinates
m WhitePlastic =
    rd .7
    rs .02 0
    sides 2
o VenetianBlinds
xf -rx -60 -a 67 -t 0 .03 0 Supports arrays
    o Slat
    v v1 =
        p -2 0 0
    v v2 =
        p 2 0 0
    v v3 =
        p 2 0 .04
    v v4 =
        p -2 0 .04
    f v1 v2 v3 v4
    o
xf
o
</Geometry>
```

# New **genBSDF** script

- Perl script calls **rtcontrib** to compute BTDF
  - forward transmittance only at this time
- Simple operation:
  - `genBSDF fenes1.rad > fenes1.xml`
- Default is to embed MGF geometry
  - options for number of samples, processes, etc.



# BSDFs in mkillum

Use *illum*      BSDF      Thickness (flag to use MGF)

```
#@mkillum l- c=a d=fenes.xml t=0 u=+Z
```

```
void polygon window
```

```
0
```

```
0
```

```
12
```

```
0
```

```
0
```

```
0
```

```
0
```

```
4
```

```
6.25
```

```
6.25
```

```
4
```

```
1
```

```
1
```

```
2.5
```

```
2.5
```

Up orientation

# When To Use “I+” with BTDF Data

- When the window geometry is unknown
- When there is no direct beam component
  - redirected components usually use BTDF



# Annual Simulation

- Using mkillum with BTDFs is fairly quick, but...
- Re-rendering a scene 2000+ times for each hour?
- We need something faster...
- Can we use daylight coefficients with BTDF data?

# Three Phase Method

- ❧ **Phase I:**

Use **rtcontrib** to get daylight coefficients relating sky patches to incident directions

- ❧ **Phase II:**

Use **rtcontrib** to relate exiting portal directions to desired measurement locations (e.g., image)

- ❧ **Phase III (time-step calculation):**

$\text{sky} * \text{incident} * \text{BTDF} * \text{exiting}$



# Our Matrix Equation

$$\mathbf{i} = \mathbf{VTDs}$$

where:

- i** is the desired result vector (radiances, irradiances, etc.)
- V** is the "View" matrix defining the lighting connection between results and exiting directions for a window group
- T** is the "Transmission" matrix defining the BTDF of the window group
- D** is the "Daylight" matrix defining the coefficients between incoming directions for the window group and sky patches
- s** is a vector of sky patch luminances for a particular time and date

In a more explicit form, this would be:

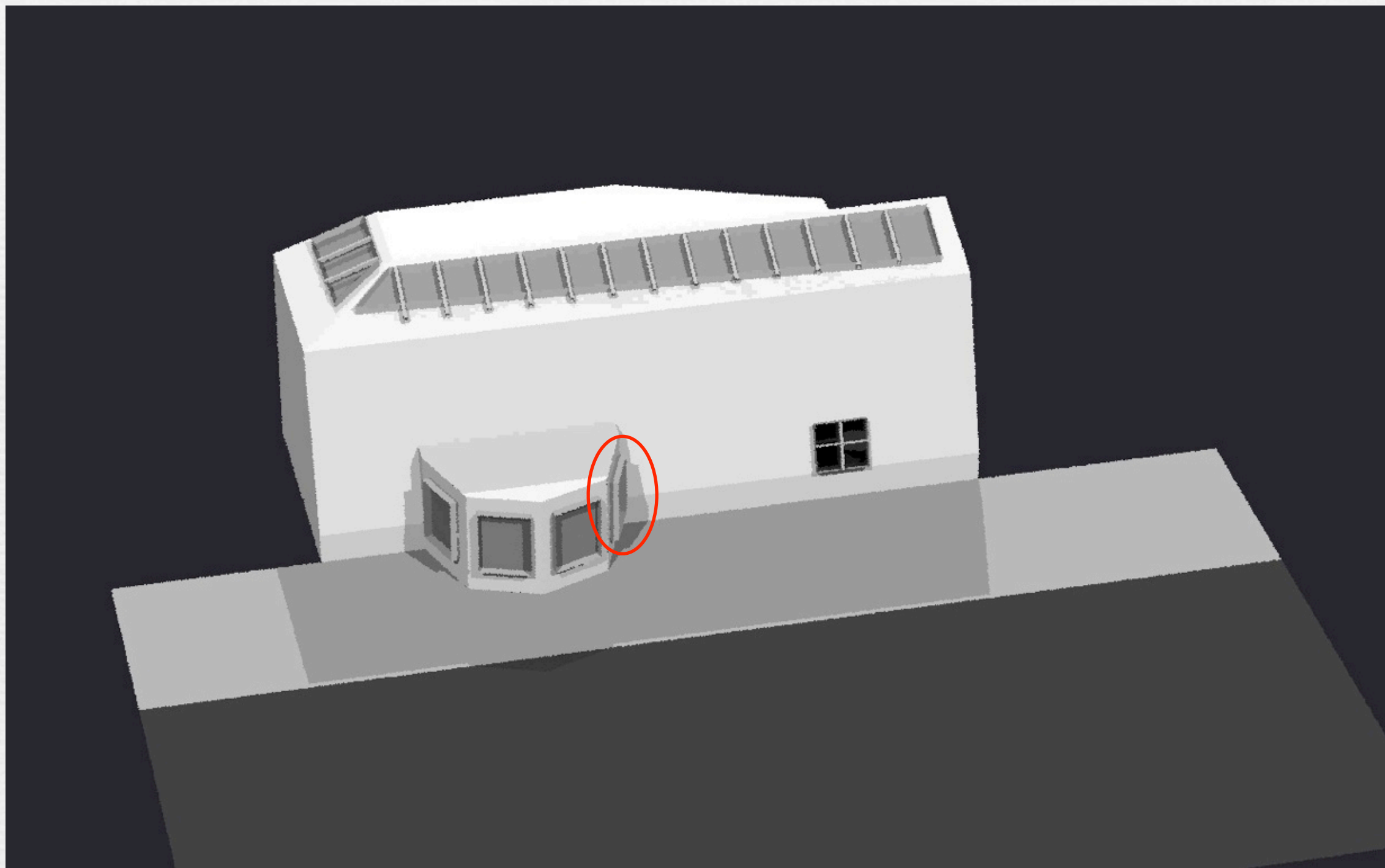
$$\begin{bmatrix} sens1 \\ \dots \\ sensM \end{bmatrix} = \begin{bmatrix} sens1edir1 & \dots & sens1edirN \\ \dots & \dots & \dots \\ sensMedir1 & \dots & sensMedirN \end{bmatrix} \begin{bmatrix} edir1idir1 & \dots & edir1idirN \\ \dots & \dots & \dots \\ edirNidir1 & \dots & edirNidirN \end{bmatrix} \begin{bmatrix} idir1dc1 & \dots & idir1dcK \\ \dots & \dots & \dots \\ idirNdc1 & \dots & idirNdcK \end{bmatrix} \begin{bmatrix} sky1 \\ \dots \\ skyK \end{bmatrix}$$

# Phase I: Compute D

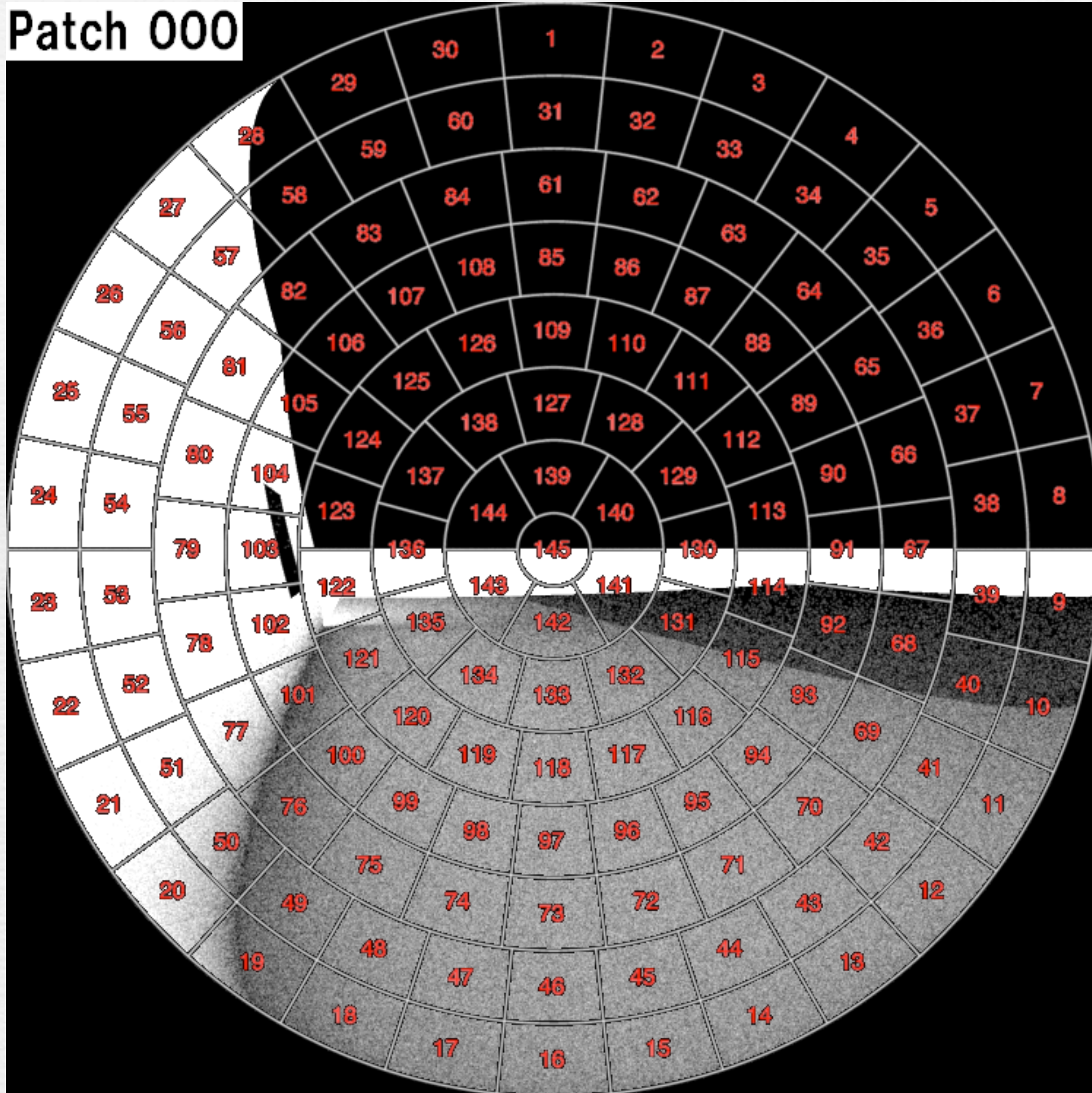
- Apply **rtcontrib** to relate sky patches to incident directions on window exterior
  - Need separate calculation for each orientation and major geometric feature
  - New **genklemsamp** utility generates samples over a given window group
    - written in Perl



# Example Space



## Patch 000





# Phase II: Compute $V$

- Use **rtcontrib** to relate sensor locations to exiting directions on window interiors
  - a single run can cover all window groups
  - **klems\_int.cal** file maps to BTDF coord.





Outgoing Directions for One Window Group



# Phase III: Time Step

- Use `genskyvec` to create sky patch vector  $\mathbf{s}$
- Use `dctimestep` to multiply it all together

Diagram illustrating the calculation of the Result Vector  $\mathbf{i}$  using the equation  $\mathbf{i} = \mathbf{VTDs}$ .

The components and their relationships are:

- BTDF** (Bidirectional Transmittance Distribution Function) points to the **T** in the equation.
- Sky Patch Vector** points to the **s** in the equation.
- View Matrix** points to the **V** in the equation.
- Daylight Matrix** points to the **D** in the equation.

The final result is the **Result Vector**  $\mathbf{i}$ .

# Phase III Example

```
gensky 9 21 12:00 -a 37.71 -o 122.21 -m 120 | genskyvec > sky.dat
pcomb '!dctimestep comp/back_SouthGroup%03d.hdr blinds1.xml SouthGroup.dmx sky.dat' \
      '!dctimestep comp/back_WestGroup%03d.hdr blinds2.xml WestGroup.dmx sky.dat' \
      '!dctimestep comp/back_NorthGroup%03d.hdr blinds2.xml NorthGroup.dmx sky.dat' \
      '!dctimestep comp/back_EastGroup_%03d.hdr blinds1.xml EastGroup.dmx sky.dat' \
> back_9-21_1200.hdr
rm sky.dat
```



# Phase III Example

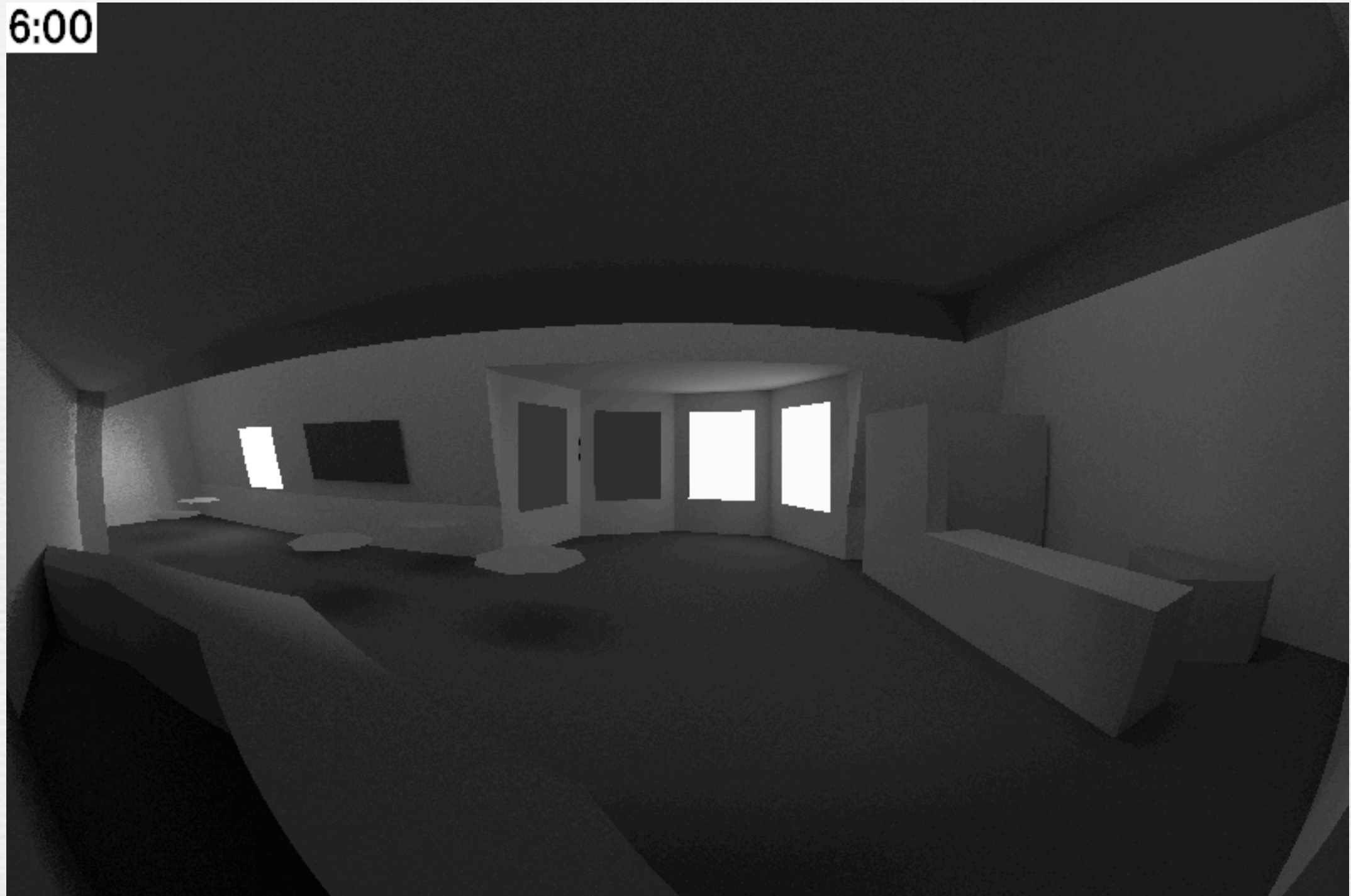
Generate sky vector for noon at the Autumn equinox

```
gensky 9 21 12:00 -a 37.71 -o 122.21 -m 120 | genskyvec > sky.dat
pcomb '!dctimestep comp/back_SouthGroup%03d.hdr blinds1.xml SouthGroup.dmx sky.dat' \
      '!dctimestep comp/back_WestGroup%03d.hdr blinds2.xml WestGroup.dmx sky.dat' \
      '!dctimestep comp/back_NorthGroup%03d.hdr blinds2.xml NorthGroup.dmx sky.dat' \
      '!dctimestep comp/back_EastGroup_%03d.hdr blinds1.xml EastGroup.dmx sky.dat' \
> back_9-21_1200.hdr
rm sky.dat
```

Each call to `dctimestep` computes contributions of one window group

Time to run the above is less than 8 seconds on my  
laptop

6:00



Equinox Simulation



# Future Work

- Support for BRDF (reflection) data
- Further validation on appropriate use
- Variable-resolution & native BSDF types
- Documentation and tutorials
- What about a user interface?
  - always my last question, always unanswered