

# Introduction to Photon Mapping

*RADIANCE Workshop 2010 – Course Advanced Fenestration*

Roland Schregle

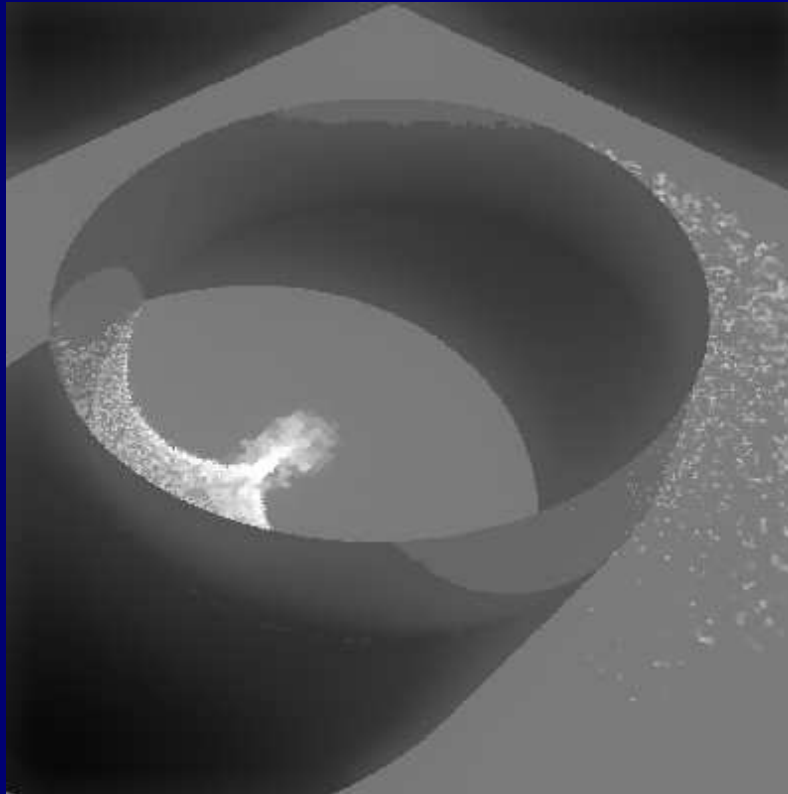
# Motivation: Caustics

Light transport from specular surfaces gives rise to *caustics* on diffuse surfaces.



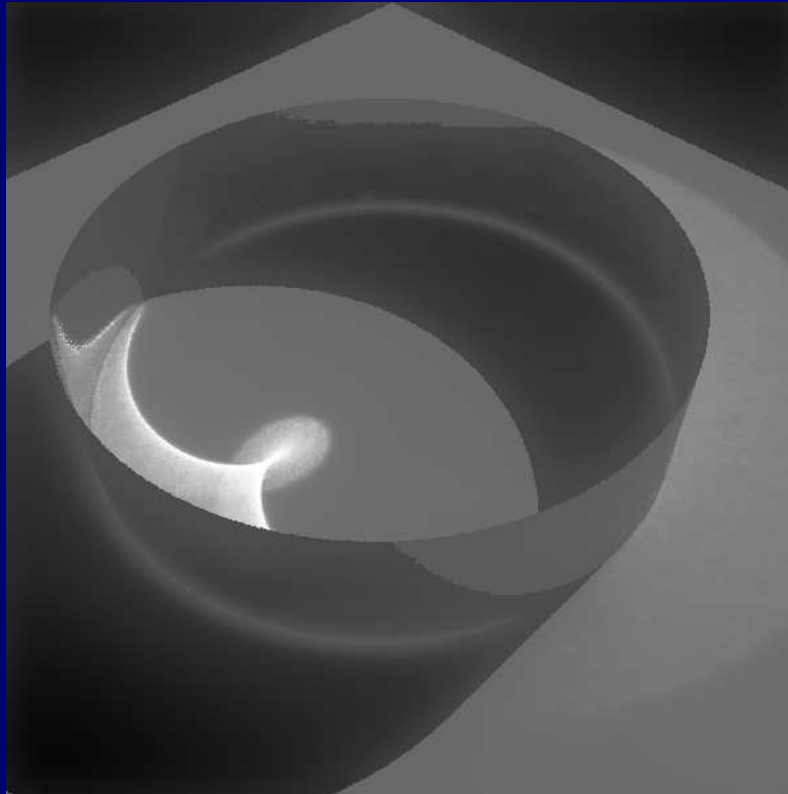
# Motivation: Caustics

RADIANCE traces rays “backwards”  
[viewer  $\rightarrow$  objects  $\rightarrow$  sources]  $\Rightarrow$  inefficient and noisy.



# Motivation: Caustics

Supplement RADIANCE with forward raytracer!  
[sources  $\rightarrow$  objects  $\leftarrow$  viewer]



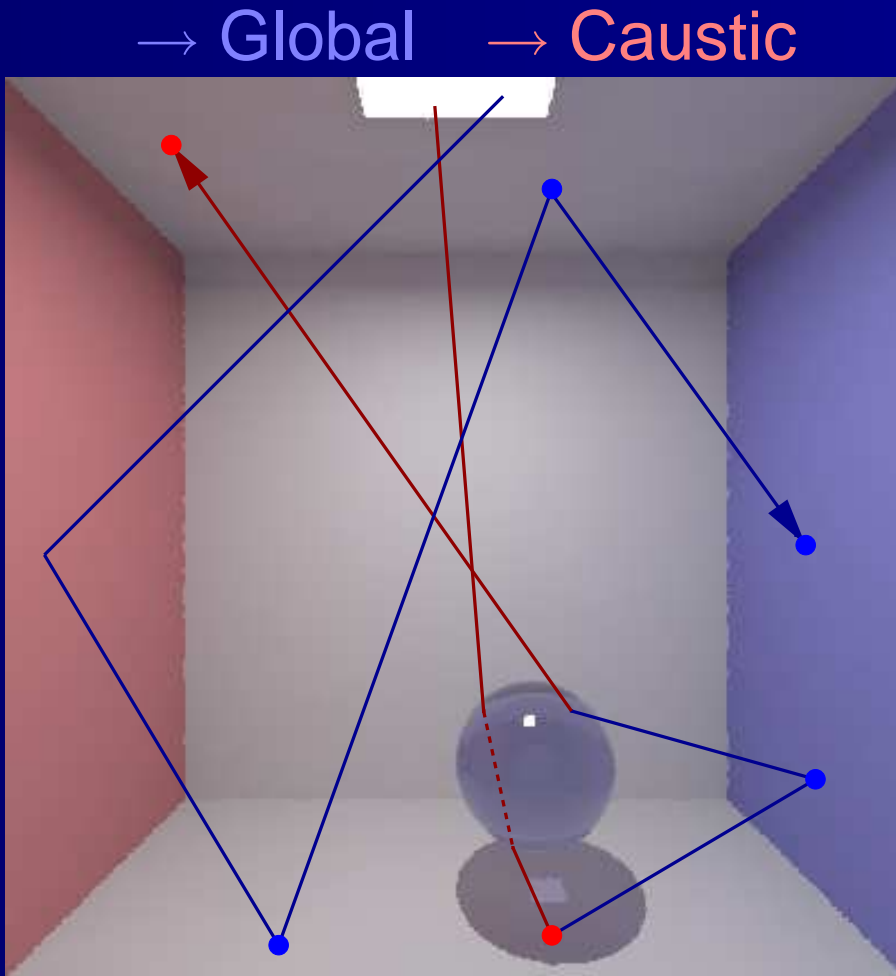
# Overview

- Two-pass method based on Monte Carlo (light) particle transport simulation [Wann Jensen 1995]  
**Forward pass** emits photons from light sources, scatters/absorbs at objects, deposits on diffuse surfaces  $\Rightarrow$  `mkpmap`.  
**Backward pass** evaluates irradiance from photons using RADIANCE's ambient calculation  $\Rightarrow$  `rpict/rtrace/rvu`.
- Photometrically validated [Schregle/Wienold 2004]

# Forward Pass

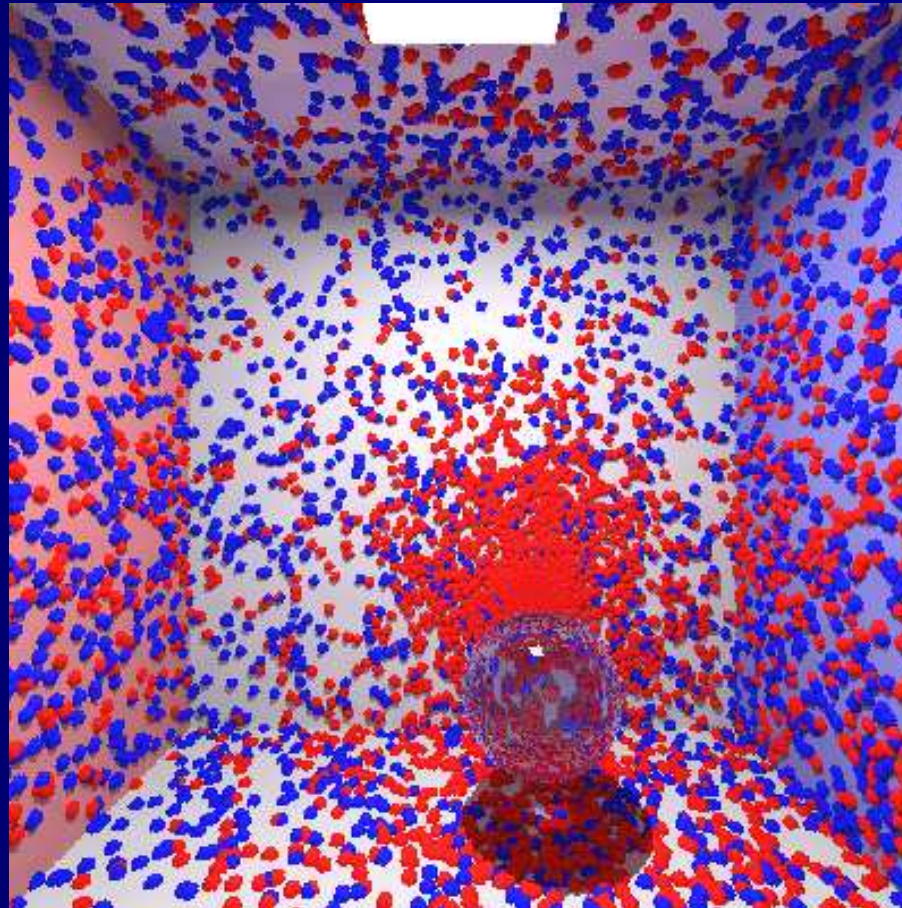
- Photons emitted from light sources, scattered at surfaces according to material properties until absorbed (*russian roulette*).
- Indirect photon hits on diffuse surfaces stored in kd-tree along with position  $\vec{x}$ , flux  $\Phi$ , normal  $\vec{N}$ .
- **global** photons stored on every indirect [diffuse|specular  $\rightarrow$  diffuse] hit.
- **caustic** photons stored on every indirect [specular  $\rightarrow$  diffuse] hit.
- **volume** photons stored in participating media (`mist`).

# Forward Pass



# Forward Pass

→ Global → Caustic





# mkpmap

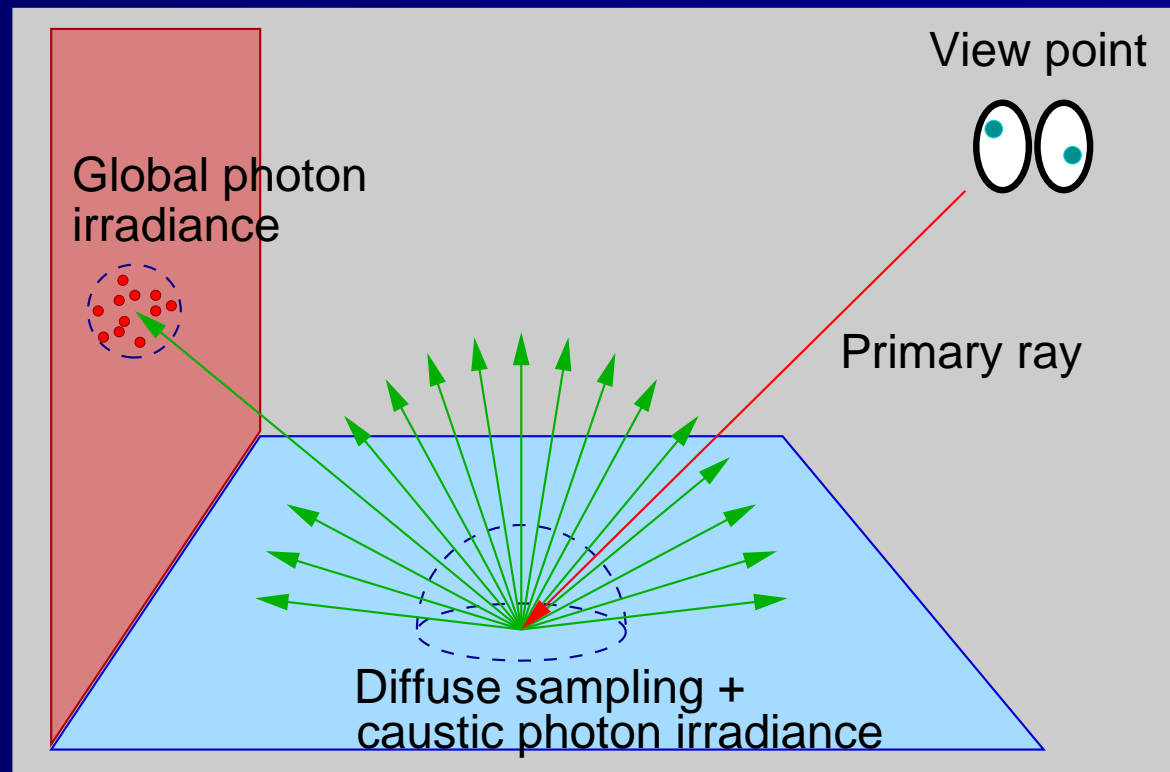
mkpmap performs forward pass and generates photon maps, e.g.

```
mkpmap -apg scene.gpm 10k -apc  
scene.cpm 50k -apv scene.vpm 100k -t  
60 scene.oct
```

generates global, caustic, and volume photon map files scene.{gpm,cpm,gpm} of approx. 10k, 50k, and 100k photons resp, with progress report every 60s.

# Backward Pass

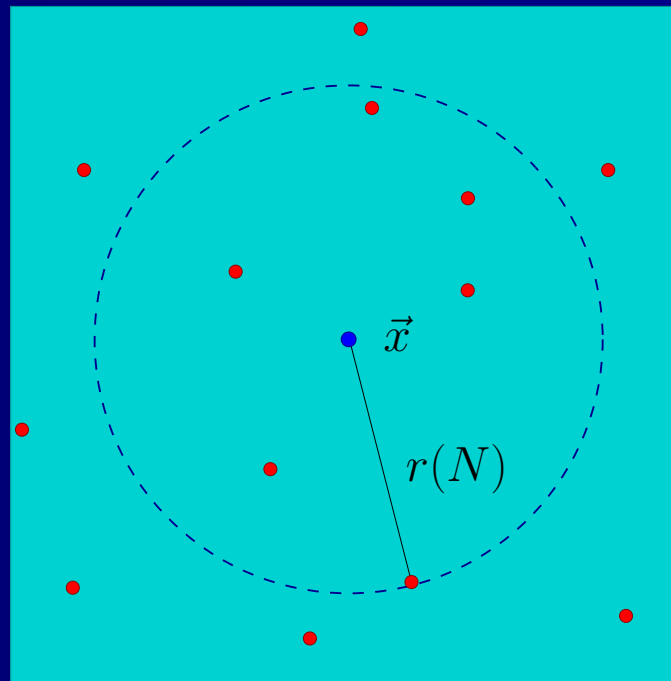
- Global photon irradiance after single ambient bounce
- Caustic photon irradiance at primary hitpoint



# Density Estimation

Irradiance  $\hat{f}$  is estimated from photon density:

- Find nearest  $N$  photons at  $\vec{x}$  in kd-tree



# Density Estimation

Irradiance  $\hat{f}$  is estimated from photon density:

- Find nearest  $N$  photons at  $\vec{x}$  in kd-tree
- Sum photon flux  $\Phi_i$  weighted by normalised kernel  $K$ :

$$\hat{f}(\vec{x}) = \sum_{i=1}^N K(\|\vec{x} - \vec{x}_i\|) \Phi_i$$

# rpict / rtrace / rvu

rpict et al. performs backward pass using photon maps, e.g.

```
rpict -ab 1 -apg scene.gpm 50 -apc  
scene.cpm 50 -apv scene.vpm 50 ...  
scene.oct
```

Loads global, caustic, and volume photon maps and computes irradiance using 50 photons per density estimate each.

Photon maps may be reused for multiple viewpoints!

# rpict / rtrace / rvu

-ab has nonstandard behaviour:

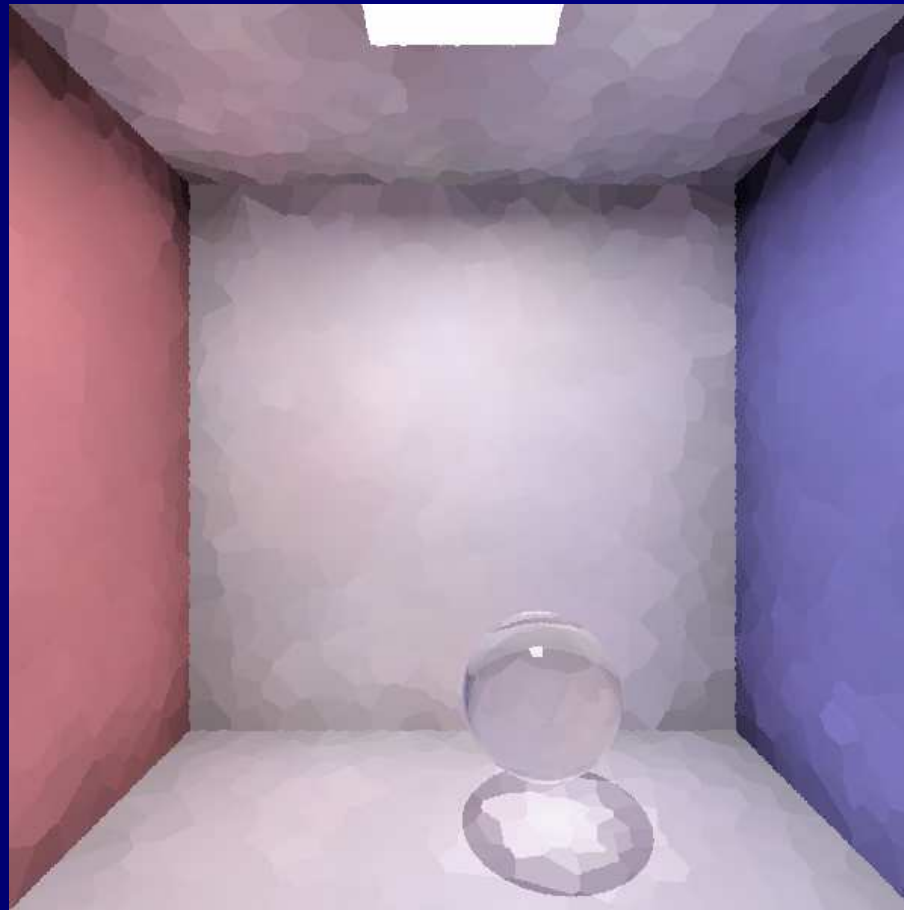
- ab > 0: global photon irradiance via *one* ambient bounce (regardless of actual -ab)
- ab < 0: no ambient bounce, global photon irradiance at primary hitpoint
- ab = 0: no ambient component at all, caustics at primary hitpoint

# Optimisations: Final Gather

Precompute and store irradiance for a fraction of global photons after forward pass [Christensen 2000]

- **PRO:** only look up single closest photon during backward pass and use precomputed irradiance directly
- **CON:** inaccurate with nonuniform ambient illum

# Optimisations: Final Gather



Precomp. global photons rendered with  $-ab < 0$



# Optimisations: Final Gather

```
mkpmap -app scene.pgpm 40k 50 -apf  
0.25 scene.oct
```

distributes 40k global photons, precomputes irradiance for  $0.25 * 40k = 10k$  (discarding rest) using 50 photons / density estimate, and outputs them to photon map file.

```
rpict -ab 1 -app scene.pgpm ...  
scene.oct
```

renders using closest precomputed global photon

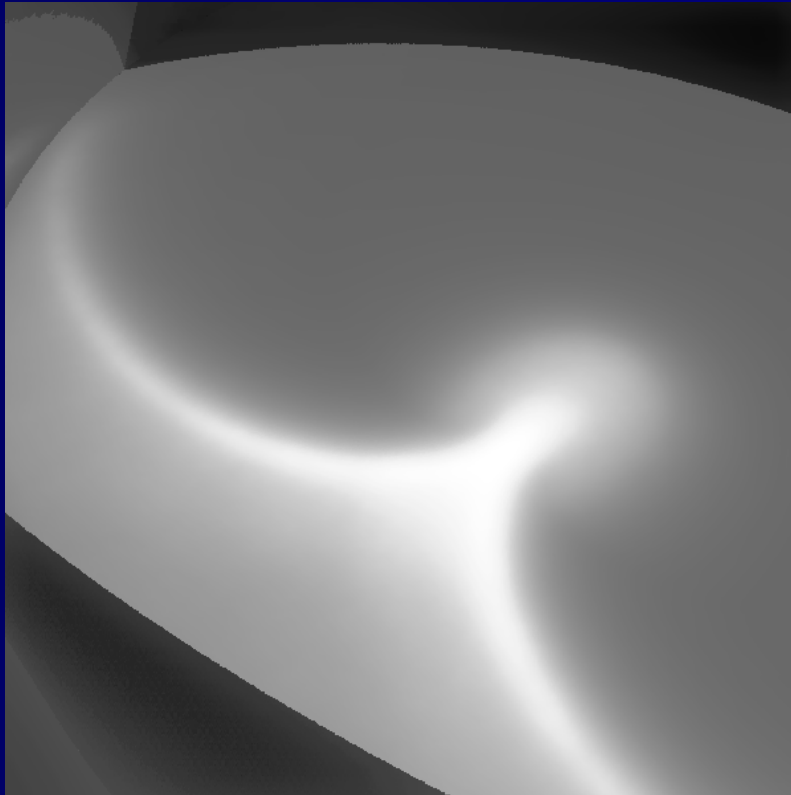
# Optimisations: Bias Compensation

Density estimation is inherently biased; photon irradiance  $\hat{f}$  converges to actual irradiance  $f$  convolved with  $K$  rather than  $f$  proper:

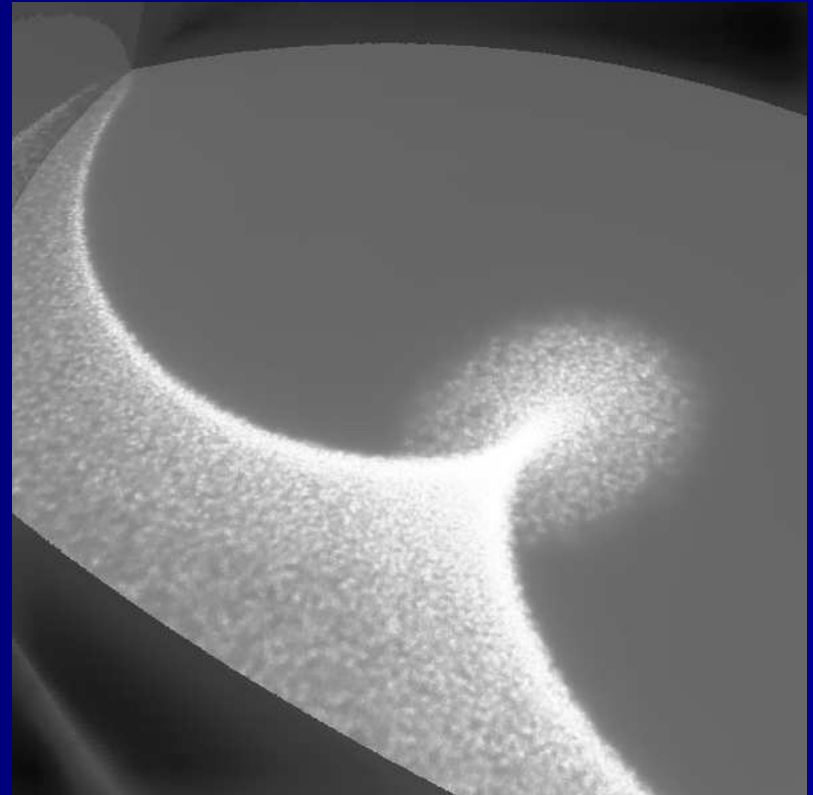
$$E \left[ \hat{f}(\vec{x}) \right] = \int_{\{\vec{y}: \|\vec{x} - \vec{y}\| \leq r(N)\}} K(\|\vec{x} - \vec{y}\|) f(\vec{y}) d\vec{y}$$

Bias visible as blur, esp. in caustics and with large  $N$   
**BUT...** reducing  $N$  increases noise!

# Optimisations: Bias Compensation



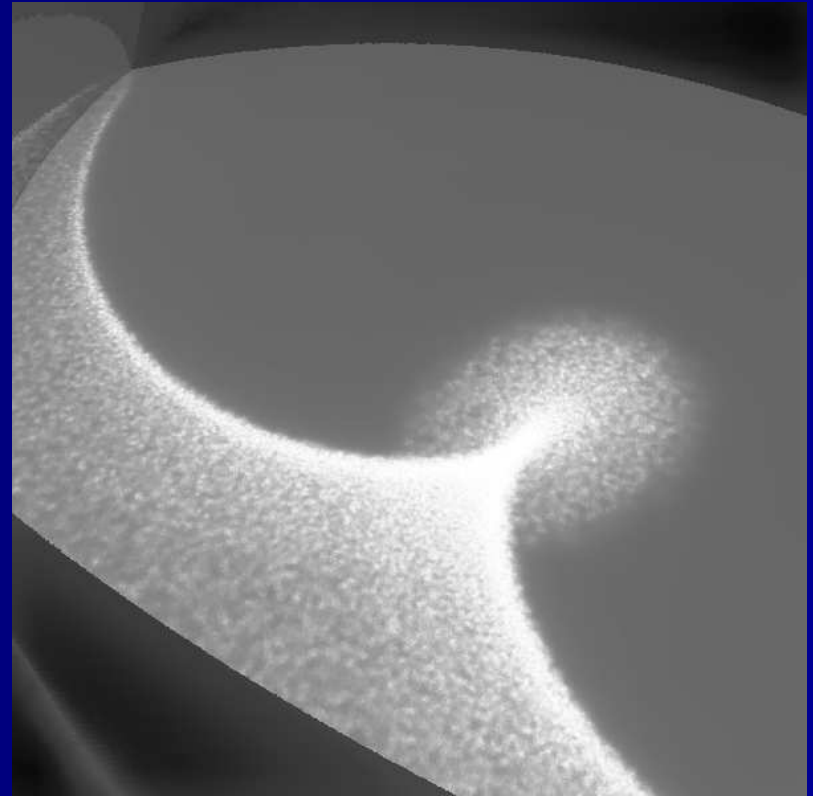
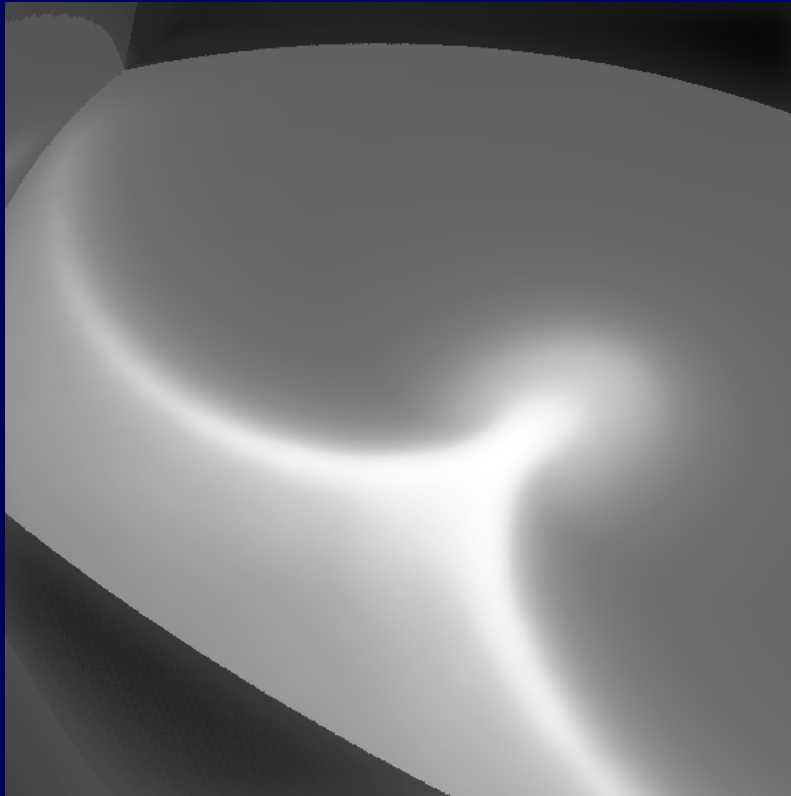
$N=2000$



$N=20$

Bias/noise tradeoff depends on photons/estimate  $N$

# Optimisations: Bias Compensation



⇒ Dynamically adapt  $N$  according to estimated bias in given interval  $[N_{min}, N_{max}]$ !

# Optimisations: Bias Compensation

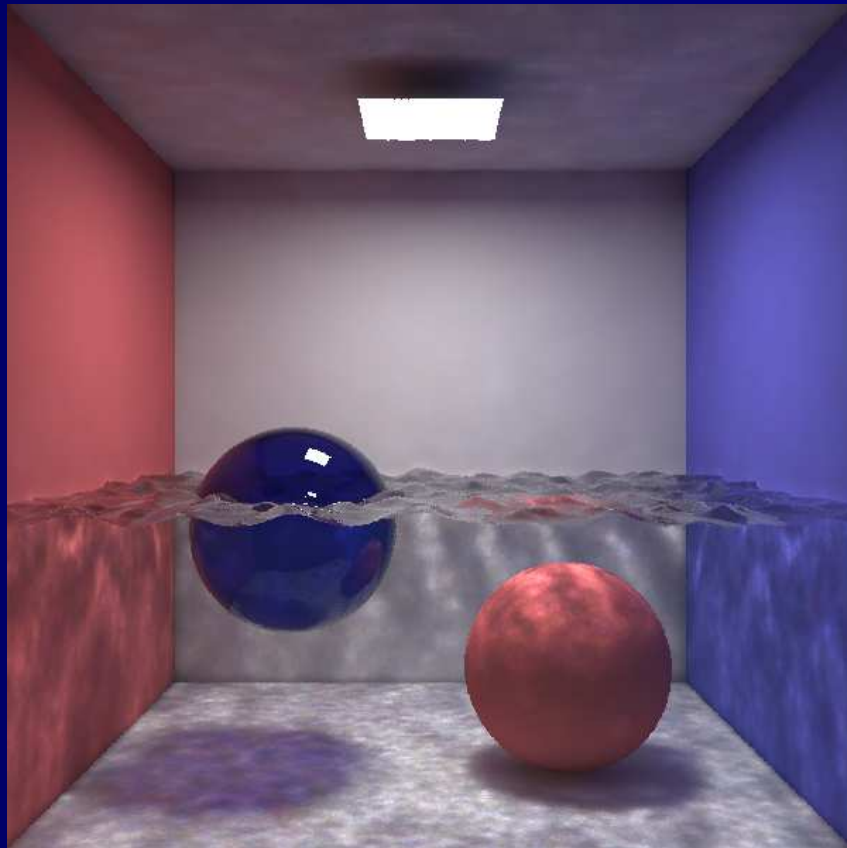
Bias Compensation adapts photons/density estimate  $N$  to minimise bias *and* noise [Schregle 2003]

- Maintains running mean and variance of density estimates to estimate bias
- Increases  $N$  until *probable* bias is detected

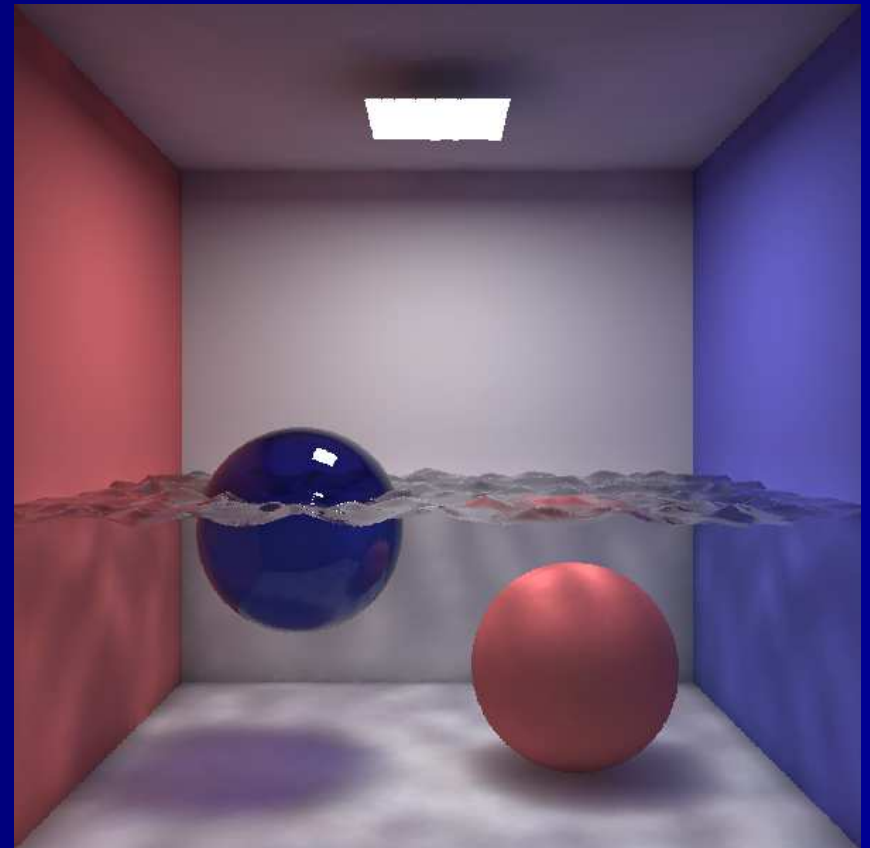
```
rpict ... -apcb scene.cpm 50 500  
... scene.oct
```

renders caustics with bias compensation using 50..500 photons/density estimate.

# Optimisations: Bias Compensation

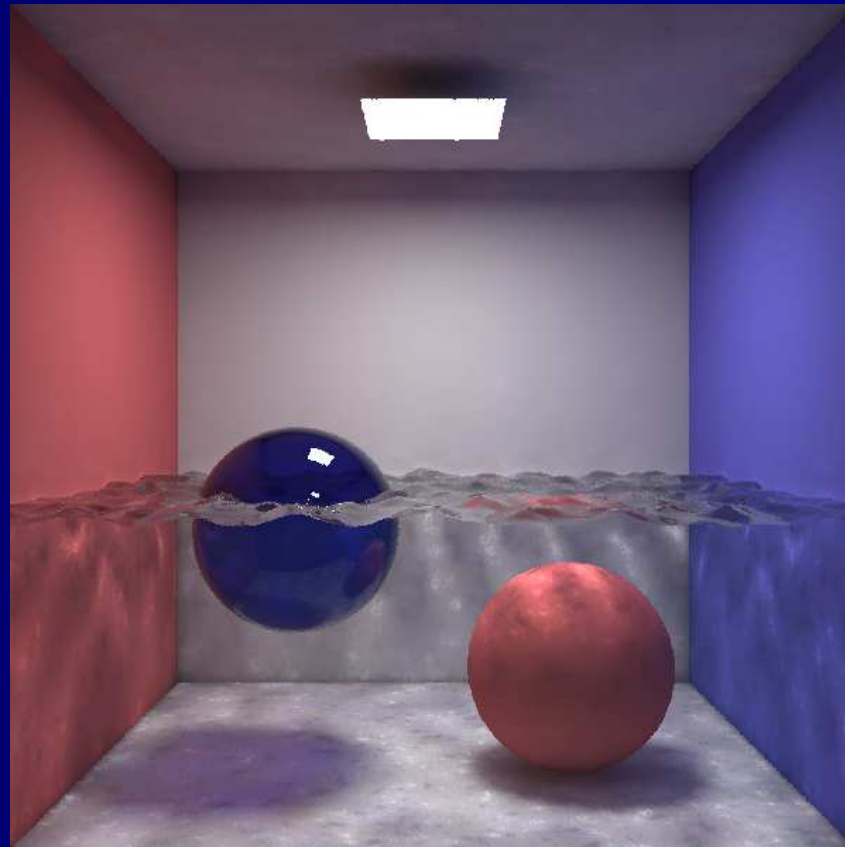


$N=50$  caustic/estimate



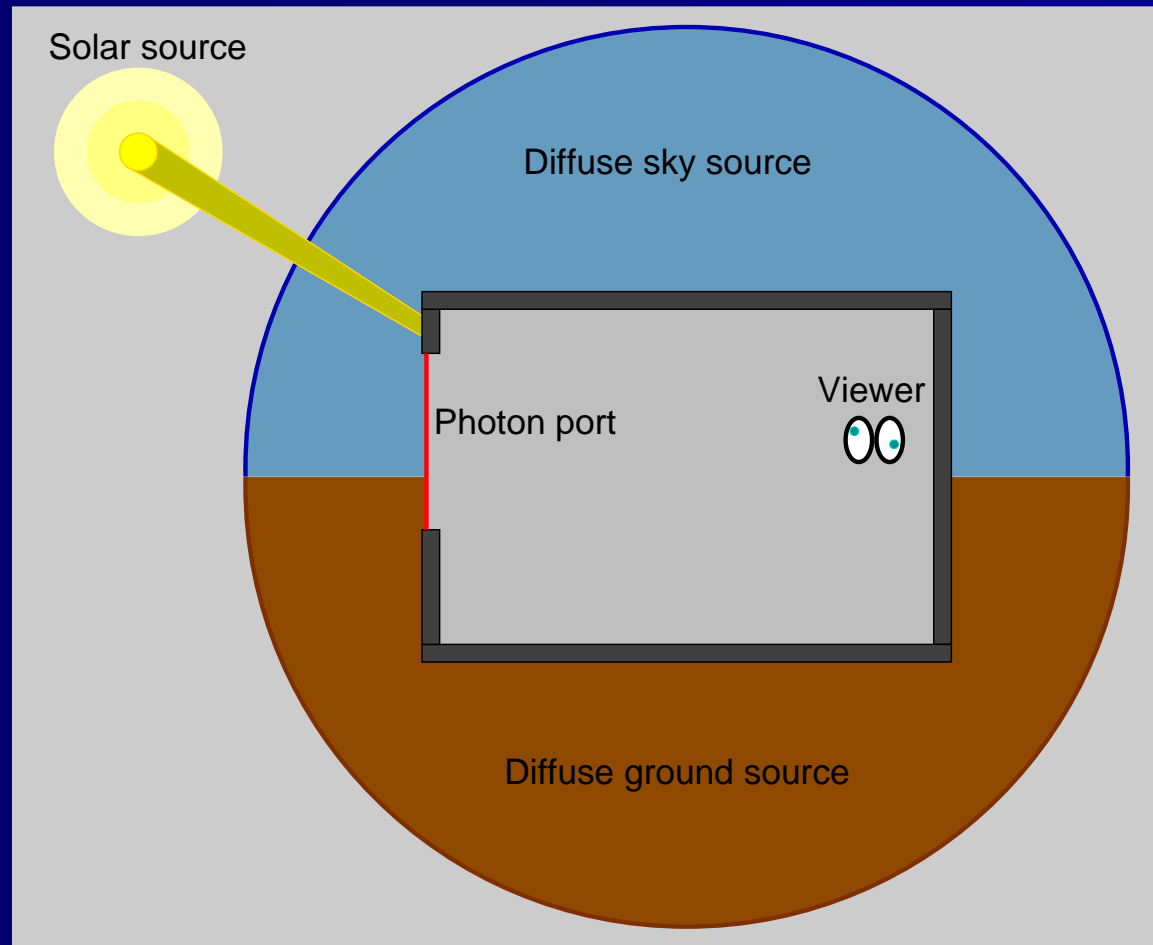
$N=500$  caustic/estimate

# Optimisations: Bias Compensation



Bias compensated, 50..500 caustic/estimate

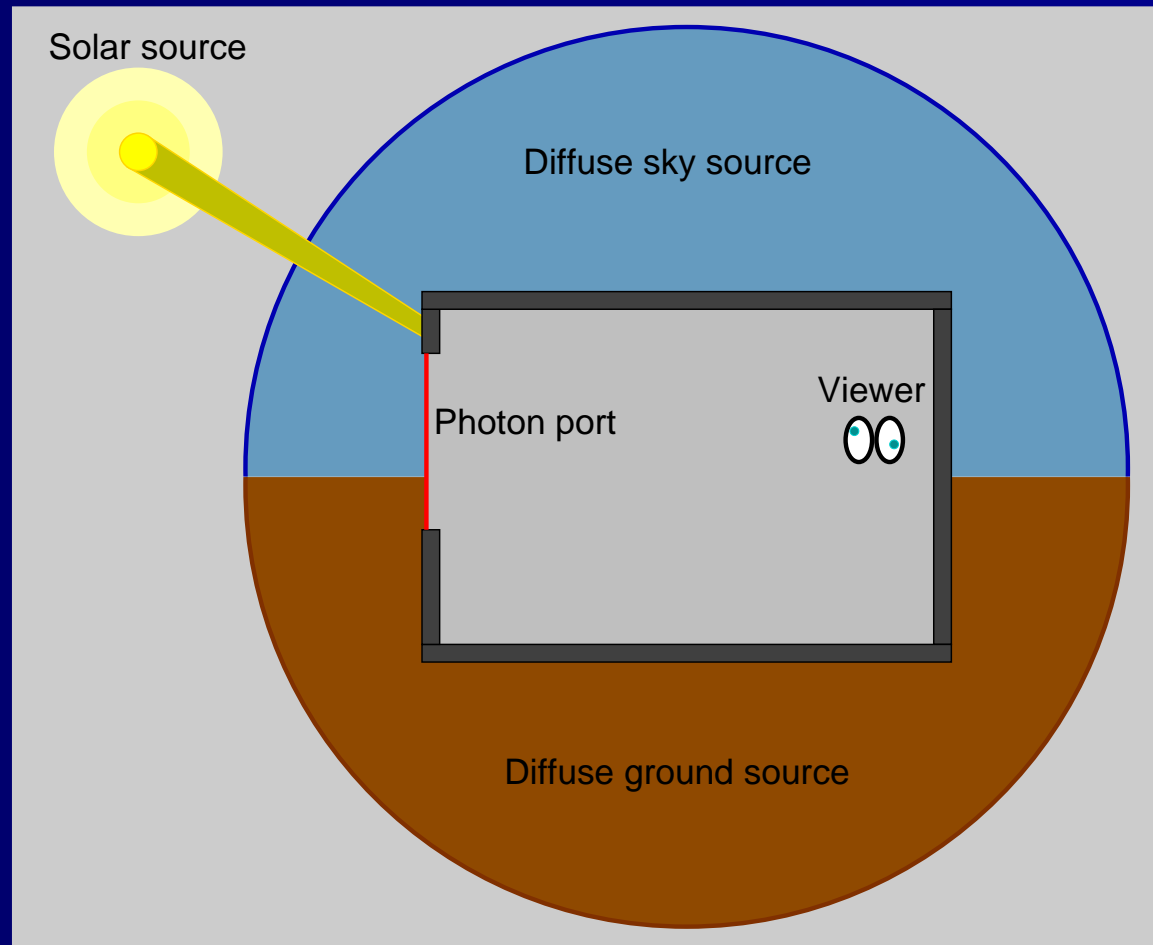
# Optimisations: Photon Ports



Few photons emitted from source reach viewer (slow)



# Optimisations: Photon Ports



⇒ Emit photons directly from window!

# Optimisations: Photon Ports

- Ports define apertures as part of regular scene geometry for photon emission from sources
- Flux incident from sources and occlusion taken into account
- Photons emitted directly from port object

Ports are specified by modifier, e.g.

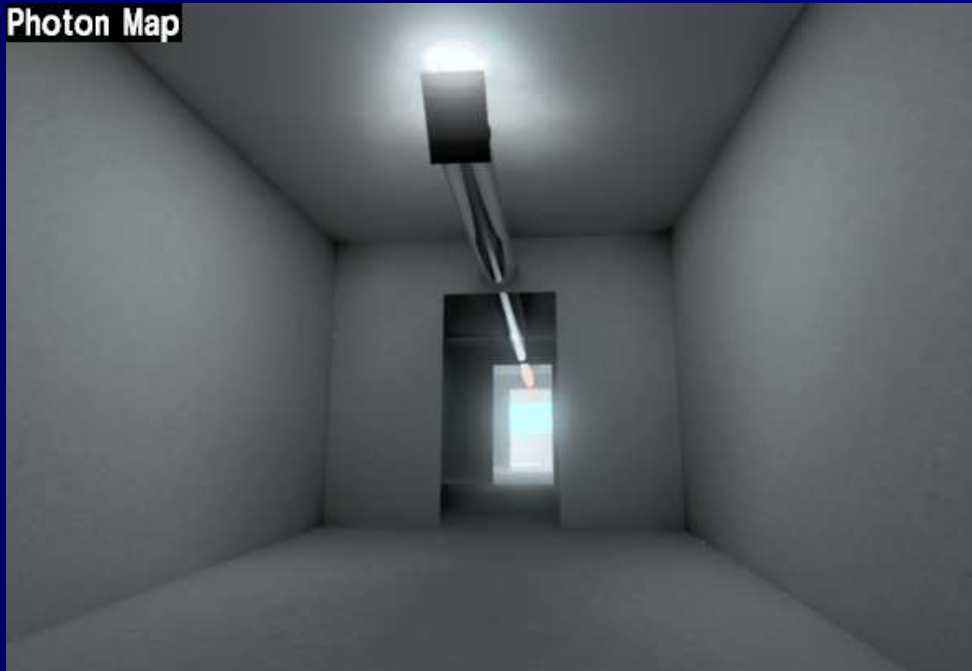
```
mkpmap -apg scene.gpm 100k -apo  
windowMat scene.oct
```

distributes photons from all objects using the windowMat material in scene.oct.

# Optimisations: Photon Ports



Photon Map



62k precomp. global  
250k caustic  
Port at pipe aperture  
(antimatter material)

# Utilities: pmapinfo

`pmapinfo <pmapFile>` gives following info on a photon map:

- photon map type (global, caustic, volume, etc)
- `mkpmap` command line used in forward pass
- Number of photons in map
- Average spectral photon flux

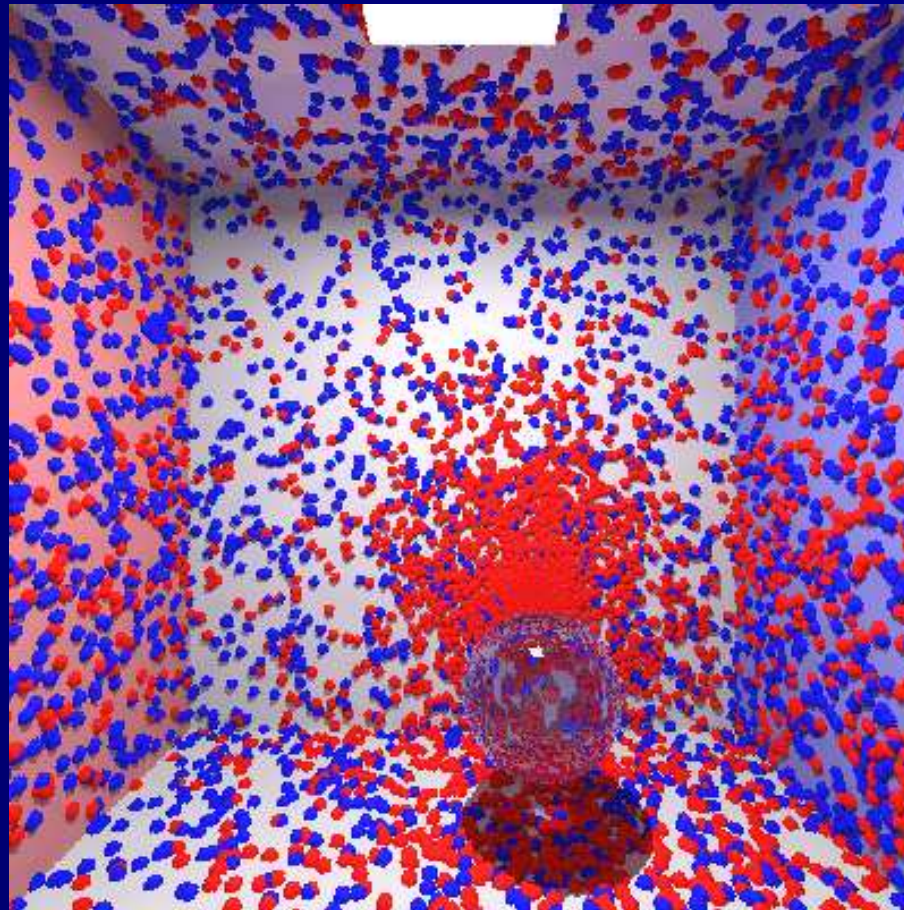
# Utilities: pmapdump

pmapdump <radius> <pmapFile>...<pmapFile>  
generates scene description of photons as spheres for  
visualisation of distribution in original scene, e.g.

```
% oconv scene.rad > scene.oct  
% mkpmap -apg pmap.gpm -apc pmap.cpm  
scene.oct  
% pmapdump 0.01 pmap.gpm pmap.cpm |  
oconv -n 16 - scene.rad >  
scene-distr.oct
```

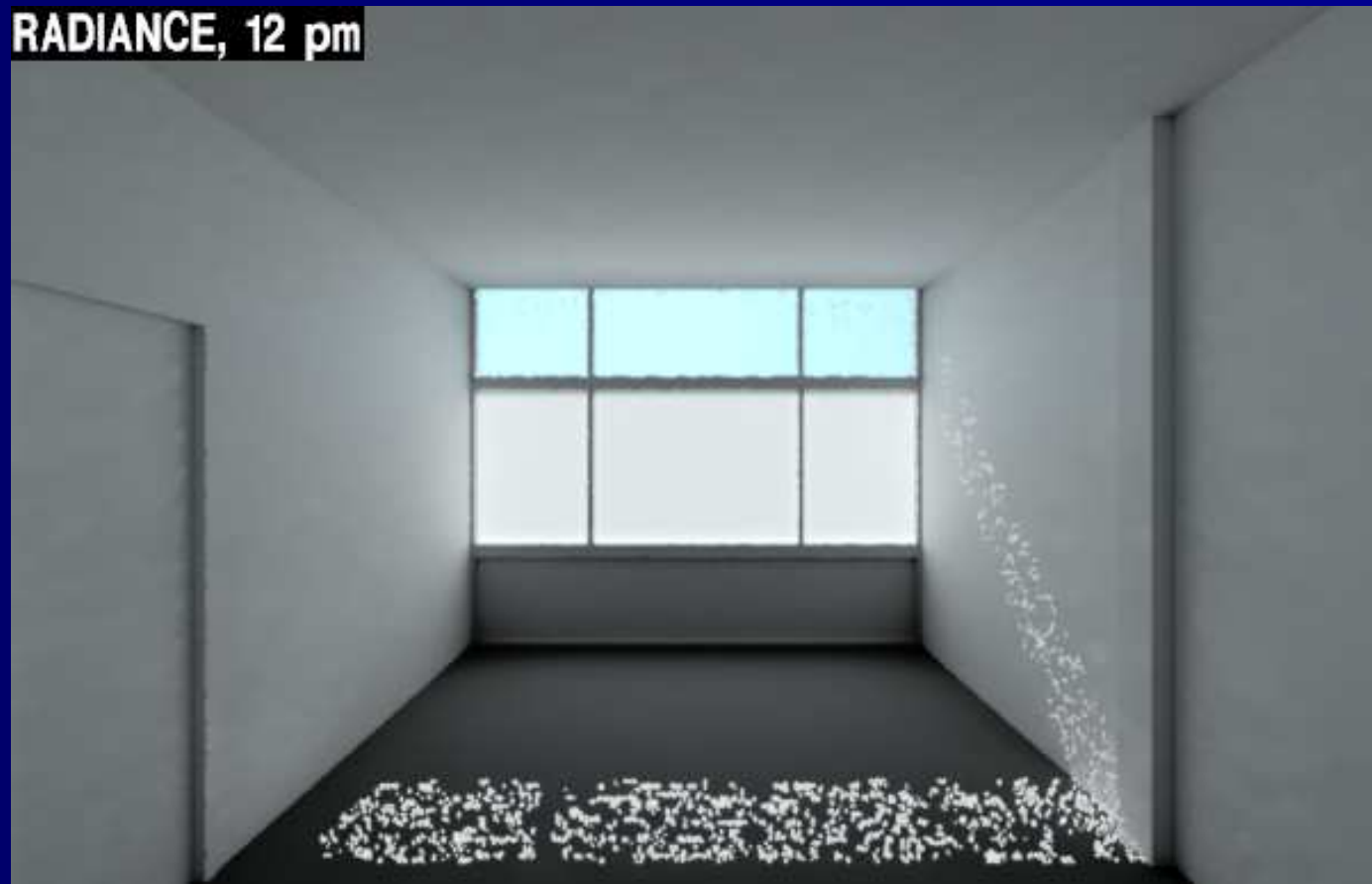
Useful for debugging with *small* number of photons.

# Utilities: pmapdump



5k global, 10k caustic dump

# Example: Y-Glass



-ab 10 -ad 4096 -aa 0.05 -ar 64

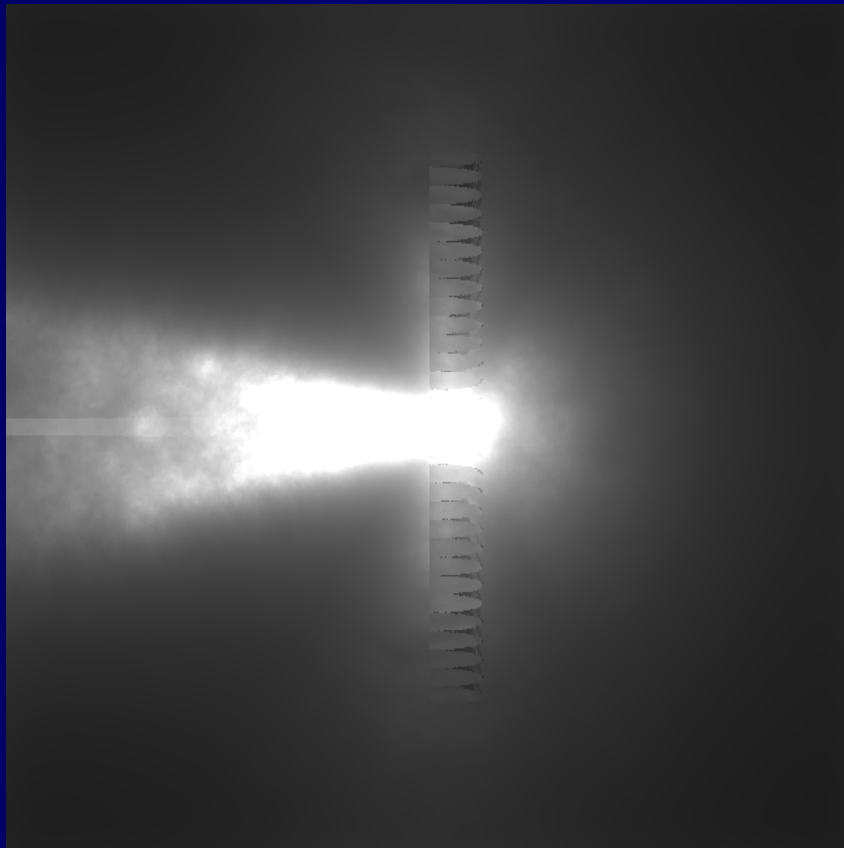
# Example: Y-Glass



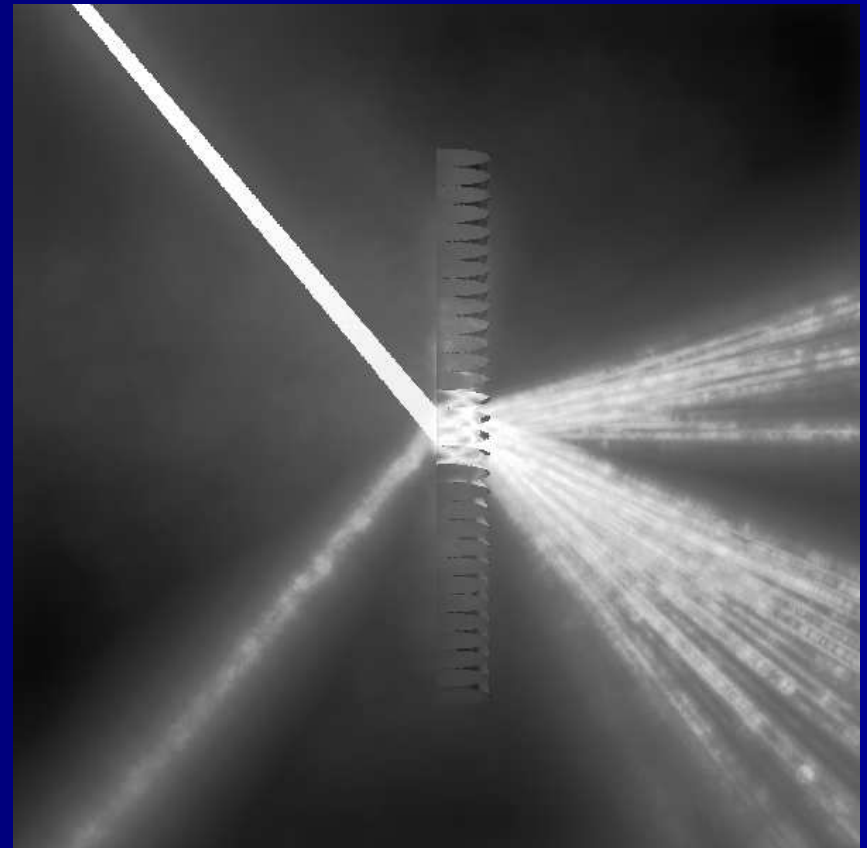
25k precomp. global, 100k caustic, ports at window



# Example: CPC

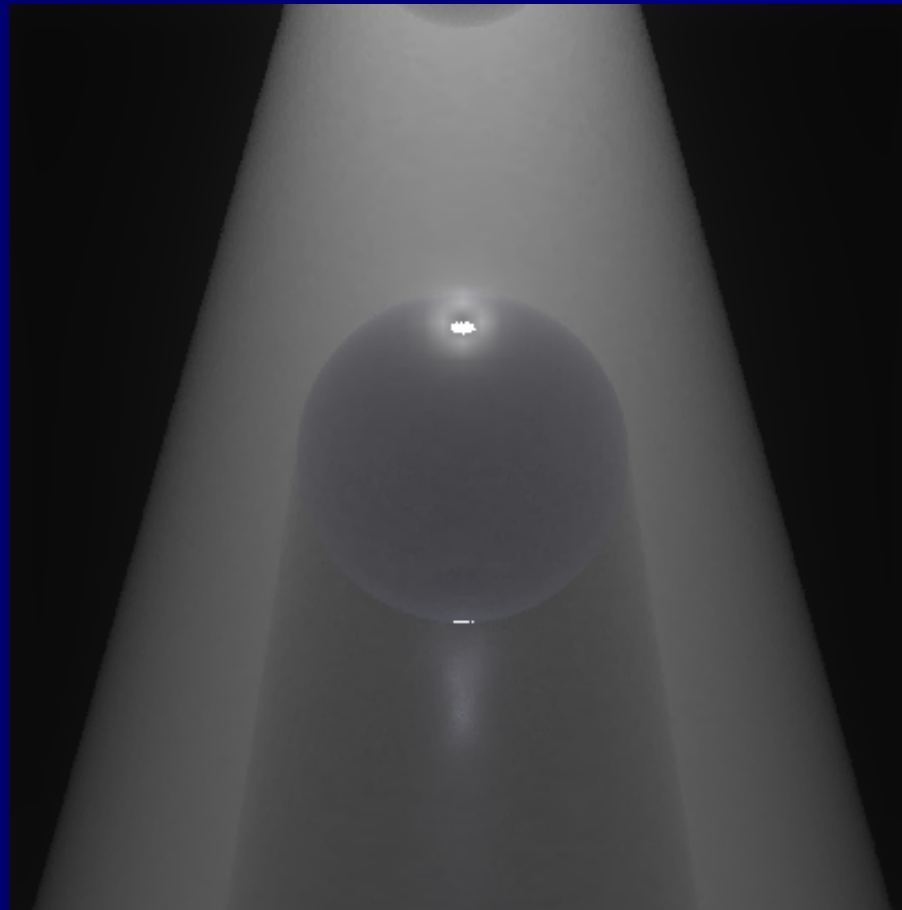


0° incidence



45° incidence

# Example: Volume Caustics



97k caustic photons

# Limitations

- Not part of official distribution – requires patching and recompiling stock RADIANCE code
- (Still) no support for user defined BRDFs (`brtdfunct`, `transfunc`, `transdata`, etc)
- Suitable parametrisation requires some experience
- Slow with complex or pathological (e.g. “leaky” or highly absorbant) scene geometry
- Inaccurate density estimates on curved surfaces or outside plane
- Bias/noise tradeoff

# Acknowledgements

**Contributors:** Peter Apian-Bennewitz, Jan Wienold,  
Christian Reetz, Carsten Bauer

**Daylight Simulation with Photon Maps:**

`www.ganjontron.net/pmap/schregle-  
daylight_simulation_with_photon_maps.pdf`

**Official website:** `www.ise.fhg.de/radiance`