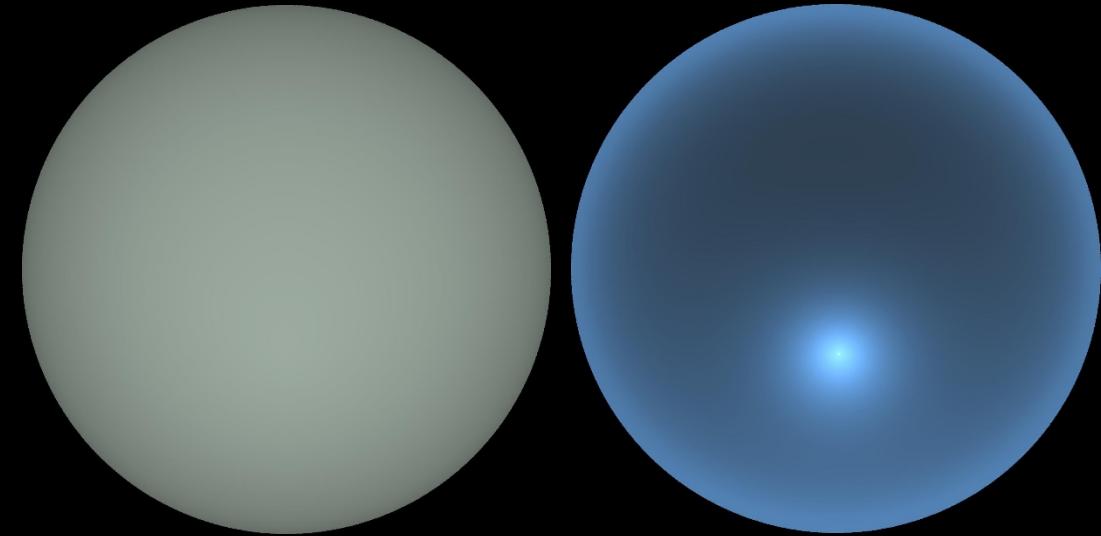


# Making simulations more **colorful**: Extension of gendaylit to create a colored sky

*J. Wienold<sup>1</sup>, A. Diakite<sup>2</sup>, M. Knoop<sup>2</sup>, M. Andersen<sup>1</sup>*



<sup>1</sup> *Ecole Polytechnique Fédérale de Lausanne EPFL*

<sup>2</sup> *Technische Universität Berlin TUB*

# Credits

All scientific results have been generated by Aicha Diakite and Martine Knoop at TU-Berlin

J. Wienold's role was the implementation into gendaylit.

# Outline

Why: *Why do we need color for the sky?*

Basis: *Spatial measurements of the sky spectrum*

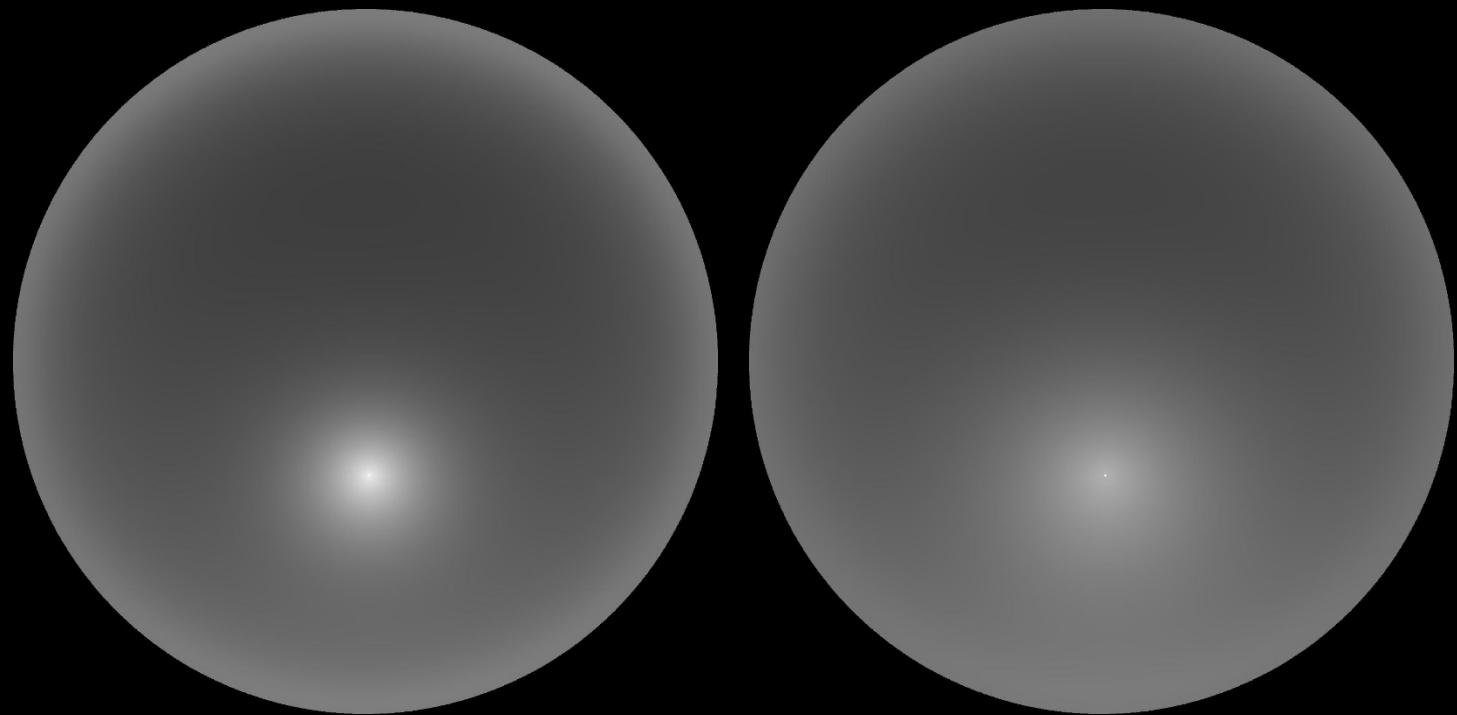
Model: *Derive of a (preliminary) color model for CIE-sky types*

Implementation: *Application of the color model onto Perez-sky distribution  
(gendaylit)*

Comparison: *Comparison of new model and Preetham*

# Why do we need a colored sky?

Status quo for daylight simulations:  
CIE and Perez Sky models are used:



No color information!

# Why do we need a colored sky?

Non-visual effects – Spectral influence - peaking in the “blue”

Visual comfort studies: To evaluate color and color rendering

Higher realism

# Measurements: Daylight Station TUB



# Spatially resolved spectral measurements



# Spatially resolved spectral measurements



- since 2014
- 145 sky patches (Tregenza)
- SPD
- 280 nm - 980 nm
- Scan every 2<sup>nd</sup> minute
- Scan duration 1 minute

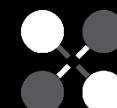
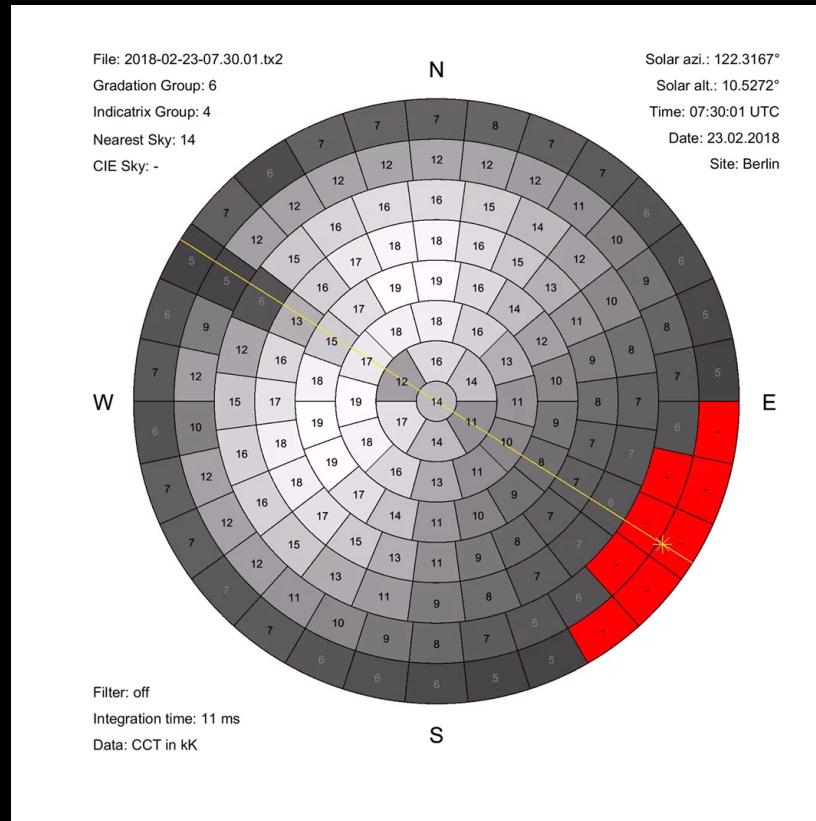
# Sky scanner for spectrally and spatially resolved measurements



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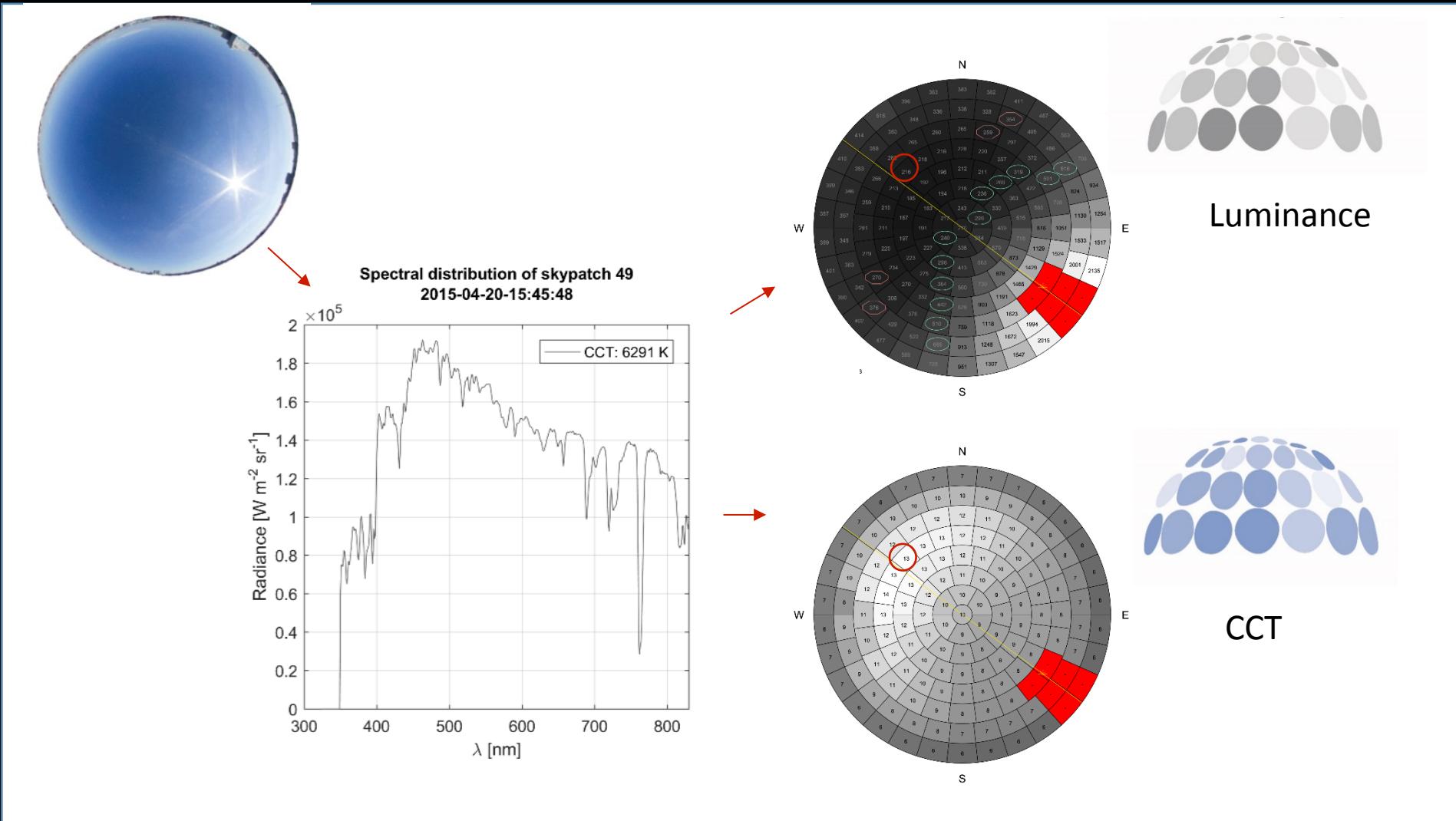
ÉCOLE POLYTECHNIQUE  
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(EPFL)

# Spatially resolved spectral measurements

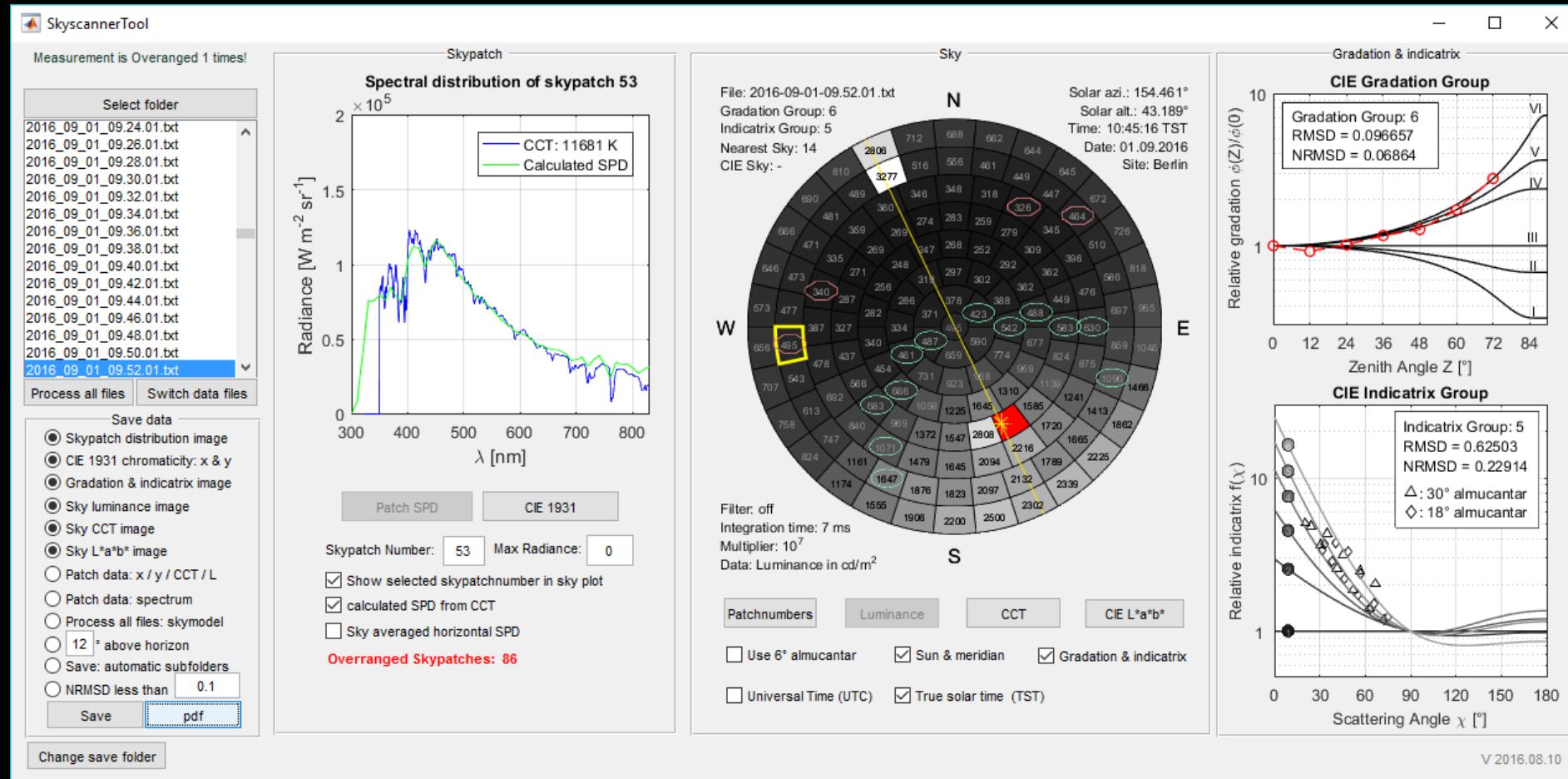


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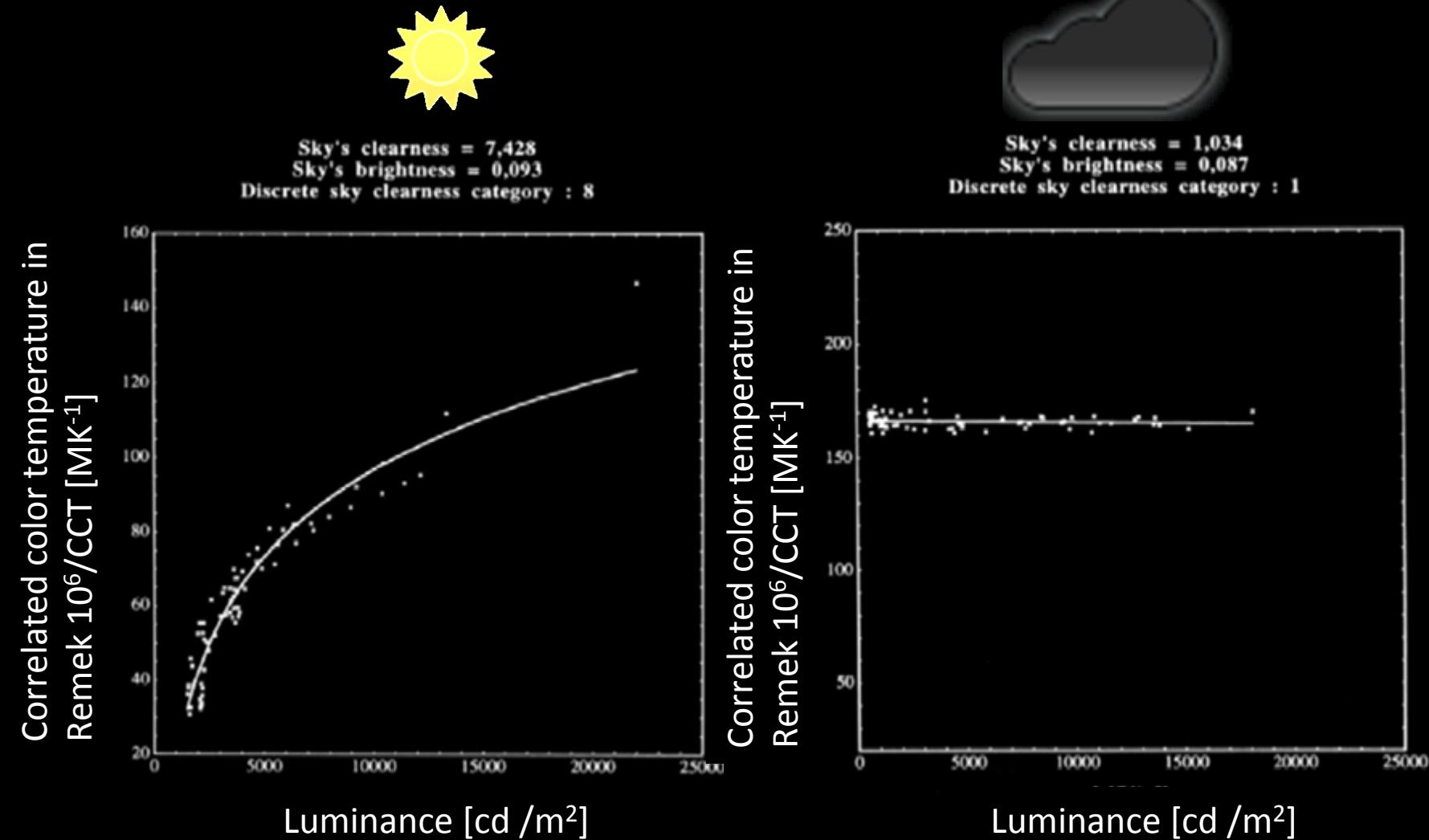
(EPFL)  
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# Data Processing: Tools



# Model



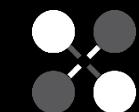
Models and Figures from:

C. CHAIN; D. DUMORTIER, M. FONTOYNONT, A COMPREHENSIVE MODEL OF LUMINANCE, CORRELATED COLOUR TEMPERATURE AND SPECTRAL DISTRIBUTION OF SKYLIGHT: COMPARISON WITH EXPERIMENTAL DATA. Solar Energy, Volume 65, Issue 5, 1 April 1999, Pages 285-295

# Model



Regression

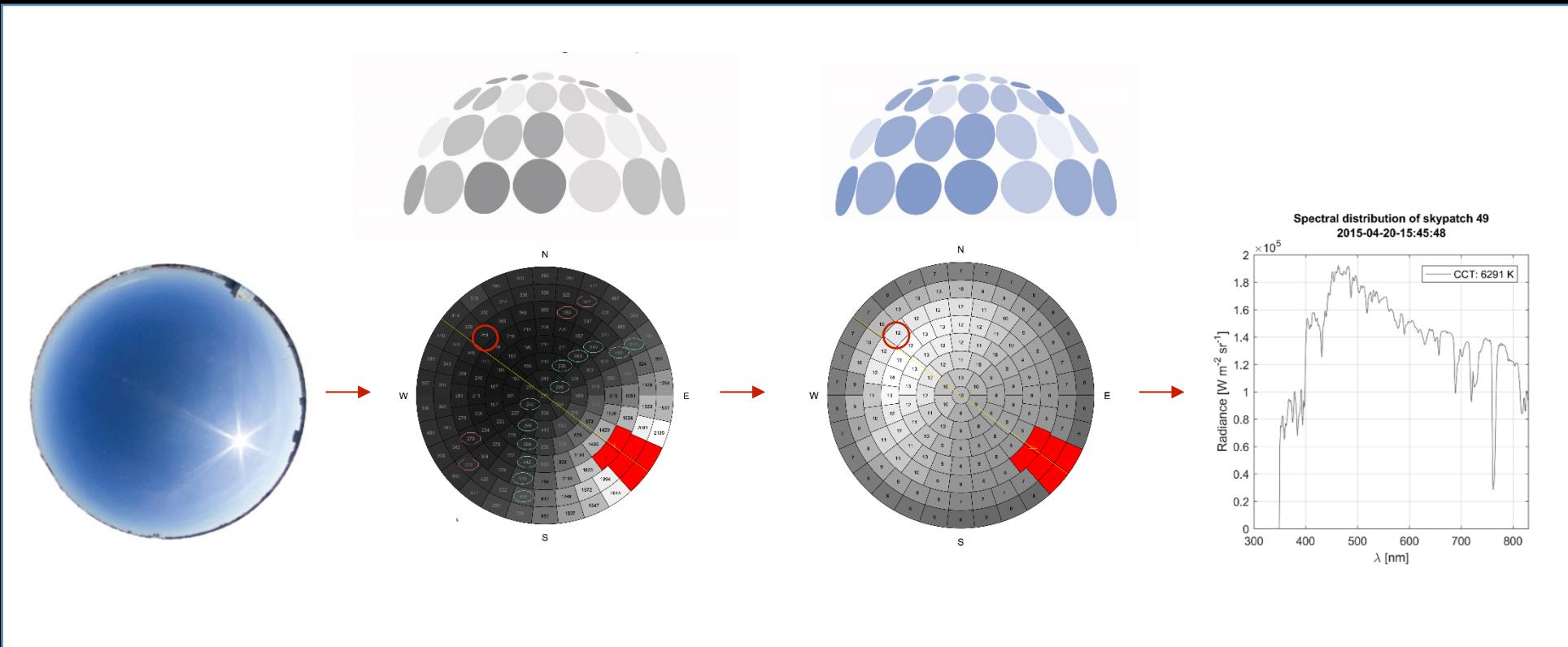


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# Implementation

Model is based on the relation

$$\text{CCT} \sim L$$

Idea:

Use existing and validated luminance sky models (CIE, Perez)

“just” change the color information

# Implementation

Luminance Model → L to CCT → CCT to x\_D, y\_D → x,y,Y to XYZ to Radiance RGB

Perez-All-Weather  
Model

$$\begin{aligned} X &= Y^*x/y \\ Y &= L \\ Z &= Y^*(1-x-y)/y \end{aligned}$$

$$\begin{aligned} R &= 2.569*X - 1.167*Y - 0.398*Z; \\ G &= -1.022*X + 1.978*Y + 0.044*Z; \\ B &= 0.075*X - 0.252*Y + 1.177*Z; \end{aligned}$$

# Implementation

Gendaylit generates modified output, colorfunc instead of brightfunc

New option(s) for gendaylit:

**-C** generate colored sky description (Default daylight locus: Berlin)

**-CIE** apply CIE daylight locus

Advanced option:

**-l**  $l_1 l_2 l_3 l_4 l_5 l_6 l_7$  user defined daylight locus, according to :

+

# Implementation - example

```
# gendaylit 6 20 12.2 -a 52.514861 -o -13.326908 -m -15 -C -W 829.5 55.15
# Local solar time: 12.07
# Solar altitude and azimuth: 60.9 1.9
void light solar
0
0
3 7.078e+06 7.078e+06 7.078e+06

solar source sun
0
0
0
4 -0.016475 -0.485873 0.873874 0.533000

void colorfunc skyfunc
4 skybright_r skybright_g skybright_b perezlum_c.cal
0
22 1.927e+01 2.896e+01 -1.006332 -0.229160 22.842973 -6.374109 1.275898 -0.016475 -0.485873 0.873874 12.009000 -4842999808.000000 2556800.000000
242.820007 0.232580 -4842999808.000000 2556800.000000 242.820007 0.232580 -1.284800 1.751900 -0.093786
```

# Implementation – examples and comparison

Summer sun: June 20, 12:12 Berlin.  $I_{\text{dir}} = 829.5 \text{ W/m}^2$

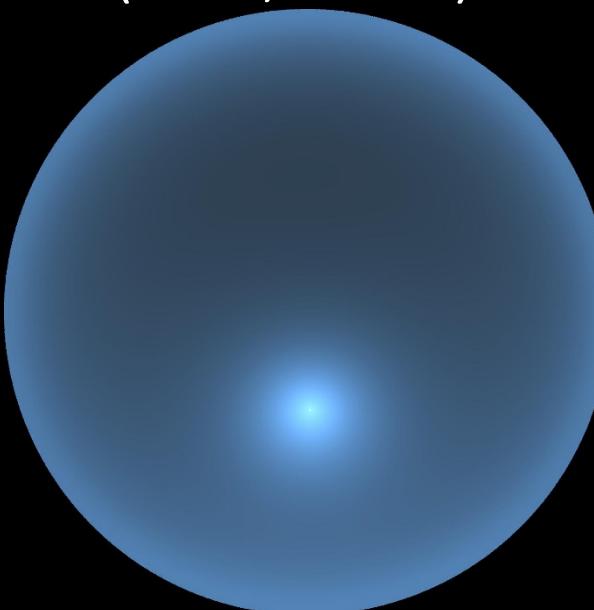
$\gamma = 60.9^\circ$ ,  $\epsilon = 12$ , very clear sky

$I_{\text{diff}} = 55 \text{ W/m}^2$

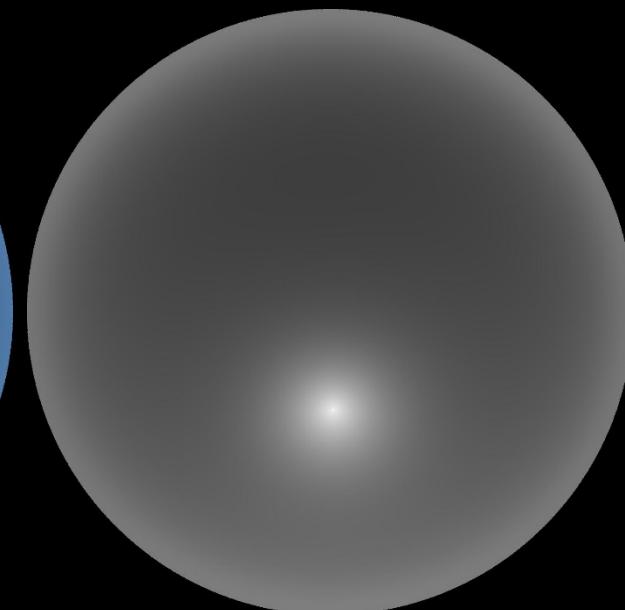
Preetham, t=2  
(M. Stock)



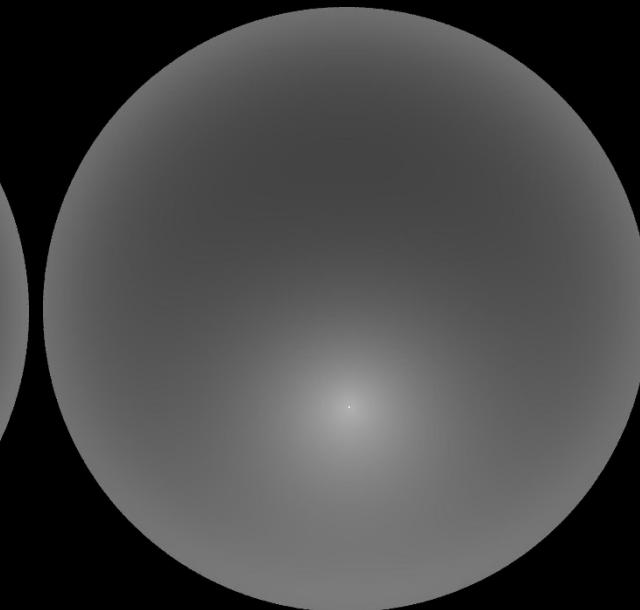
Perez  
(color , Diakite)



Perez  
(gendaylit)



CIE  
(gensky)



# Implementation – examples and comparison

Winter sun: February 18, 13:14 Berlin.  $I_{\text{dir}} = 755 \text{ W/m}^2$

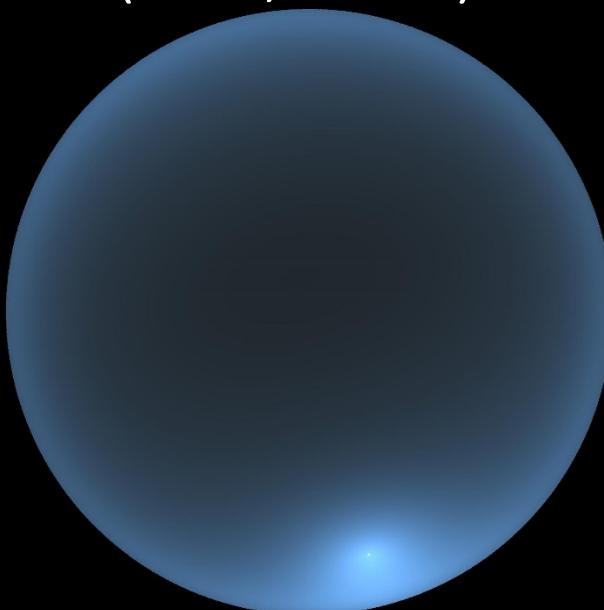
$I_{\text{diff}} = 17 \text{ W/m}^2$

$\gamma = 15.8^\circ$ ,  $\epsilon = 12$ , very clear sky

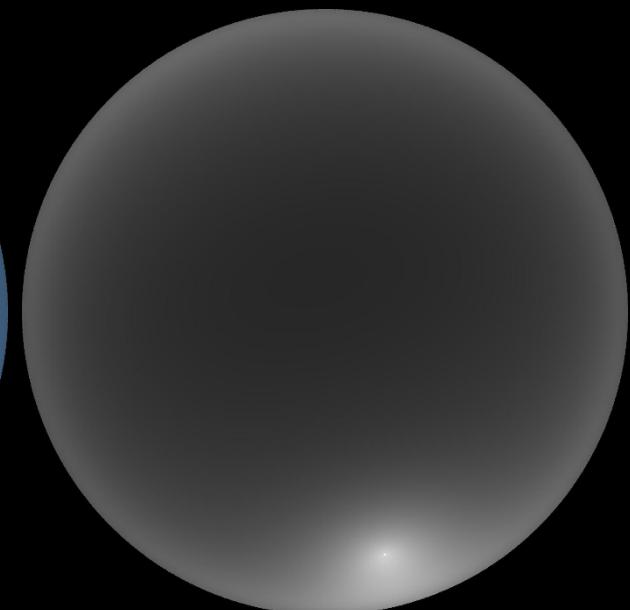
Preetham, t=2  
(M. Stock)



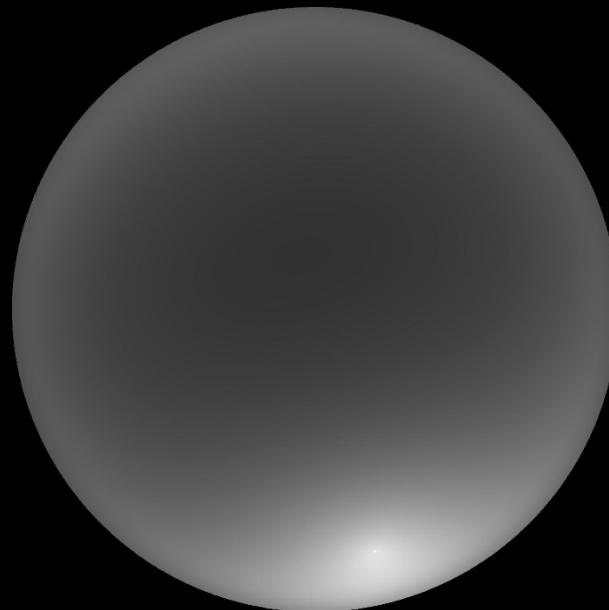
Perez  
(color , Diakite)



Perez  
(gendaylit)



CIE  
(gensky)

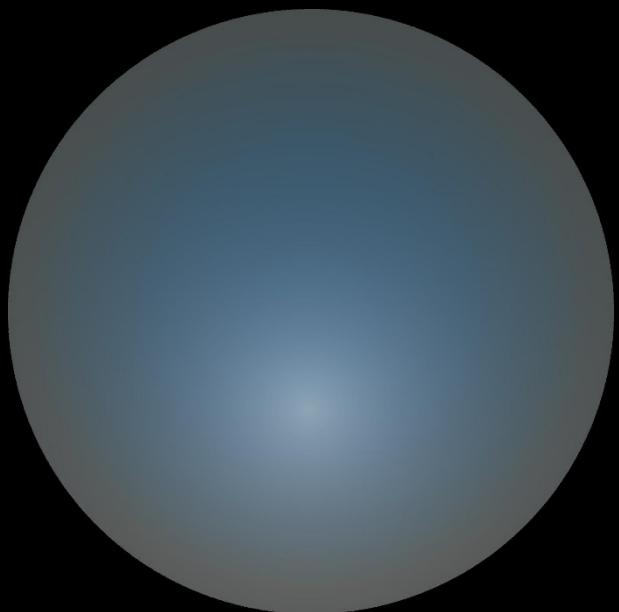


# Implementation – examples and comparison

Overcast sky: June 17, 12:04 Berlin.  $I_{\text{dir}} = 0 \text{ W/m}^2$      $I_{\text{diff}} = 193 \text{ W/m}^2$

$\epsilon = 1.00$ , overcast

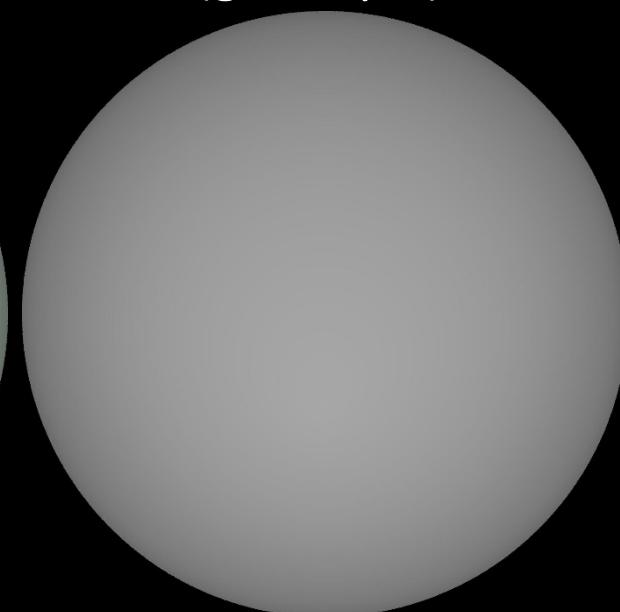
Preetham, t=30  
(M. Stock)



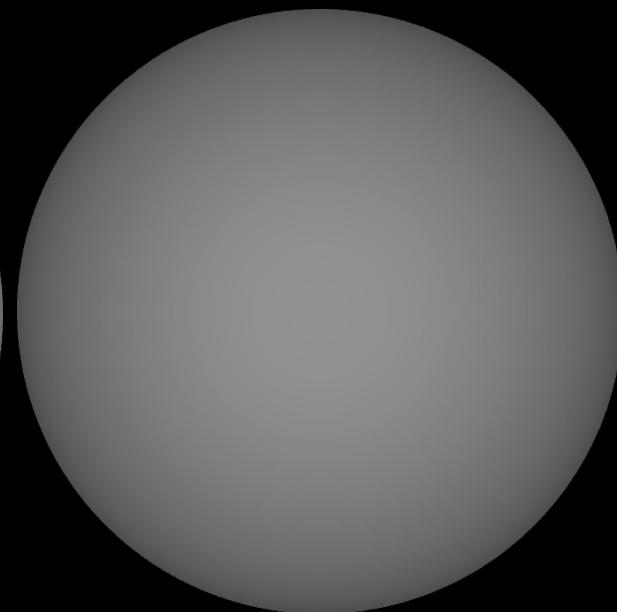
Perez  
(color , Diakite)



Perez  
(gendaylit)



CIE  
(gensky)



# Implementation – status

Preliminary model

New version “checked-in” in head release and in Radiance installers on NREL server

gendaylit is fully “downwards compatible”

At the moment: Only for clear skies  $\epsilon > 4.5$  or overcast skies  $\epsilon < 1.065$

Intermediate skies: Automatic fall-back to “grey” model

# Outlook

Intermediate skies

Full validation of the model (Development data: 1<sup>st</sup> year, validation data 2<sup>nd</sup> year)

Validation/adaptation to other locations: Large measurement campaign planned

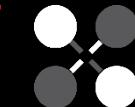
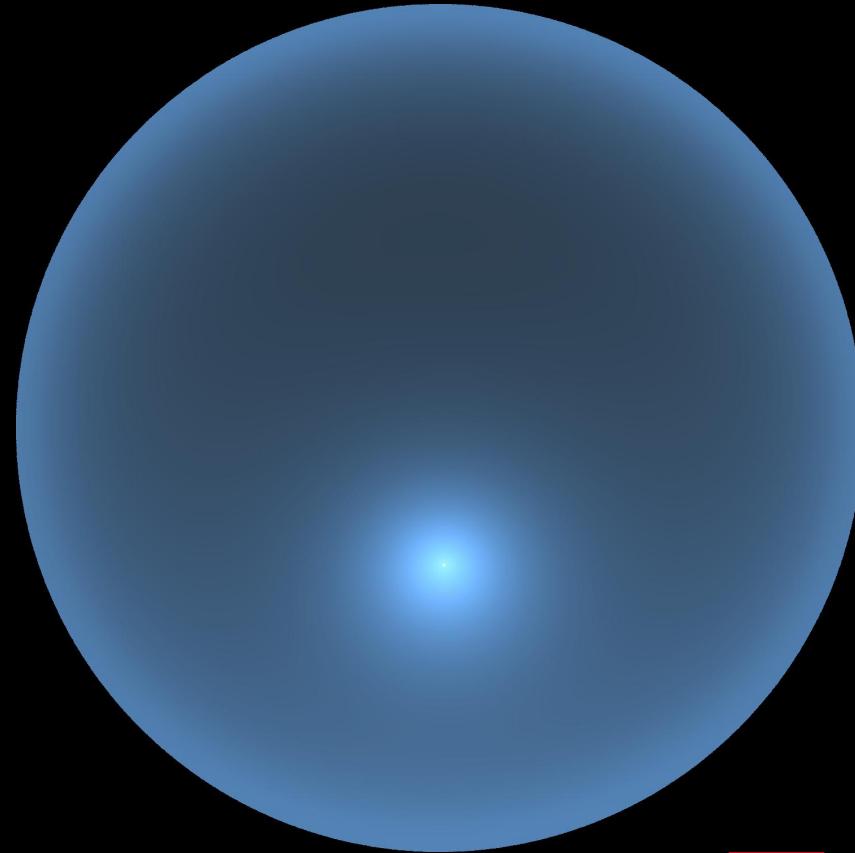
Publication

# Acknowledgement

This research is supported by the Velux foundation

# Thank you for your attention!

And have a colorful coffee break!



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